# A new South America instrumental network for monitoring of the atmospheric electric field

Jean-Pierre Raulin

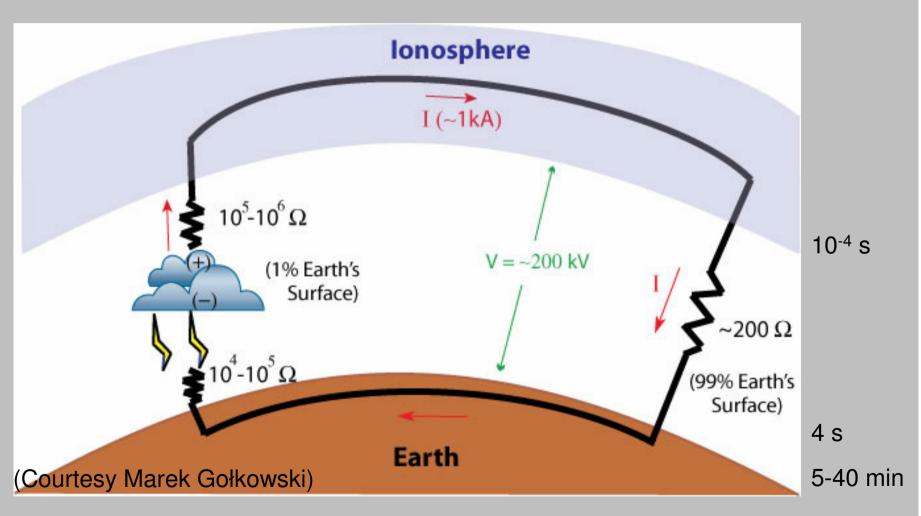


Centro de Radioastronomia e Astrofísica Mackenzie, Universidade Presbiteriana Mackenzie, Escola de Engenharia, São Paulo, SP, Brazil

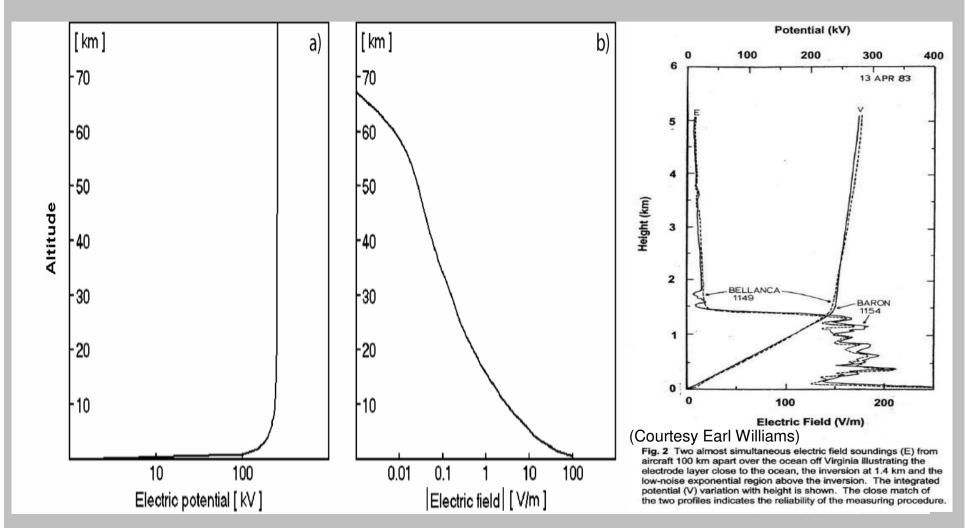
E. Macotela, R. Escate, Christian Ferradas, F. Bertoni, E. Correia, P. Kaufmann

# The Global Atmospheric Electric Circuit

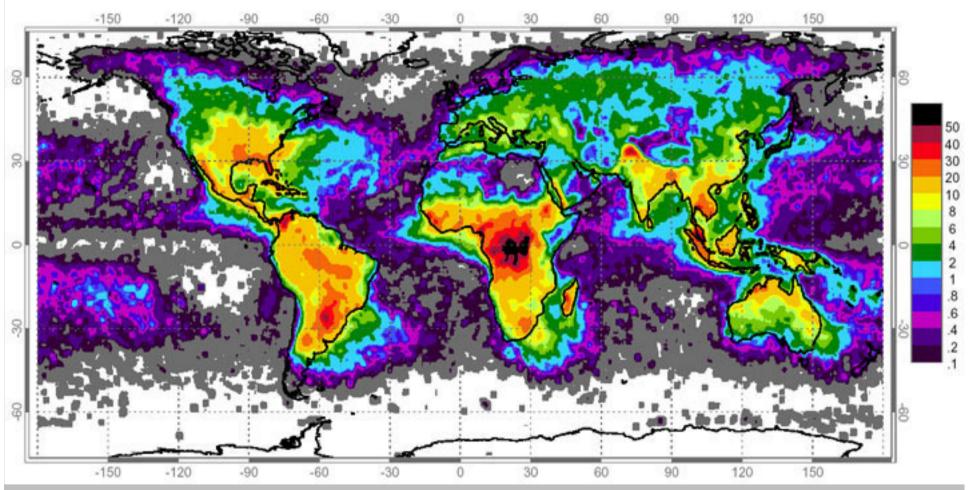
The global atmospheric electrical circuit links charge separation in disturbed weather regions with current flow in fair weather regions



# Atmospheric electric parameters, fair weather Model results of Rycroft et al., JASTP, **62**, 2485, 2007

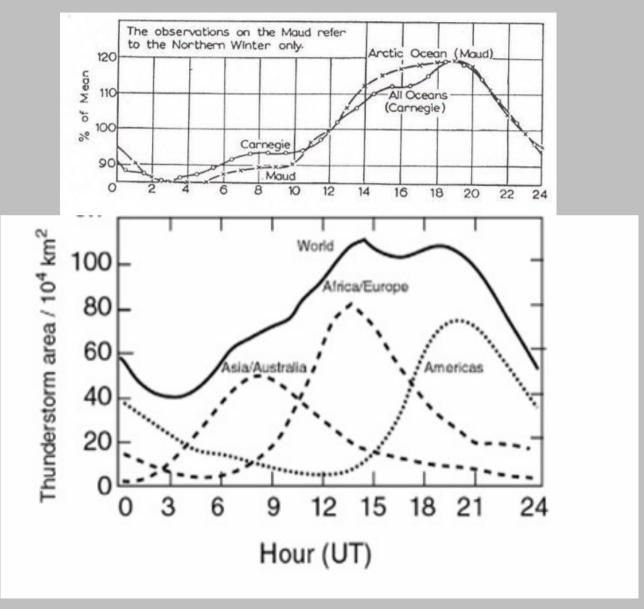


Global distribution of the number of lightning flashes per square km per year, measured from space (NASA): thunderstorms, an important generator in the global circuit



3rd ISWI Workshop – 8-12/10/2012 - Quito - Ecuador

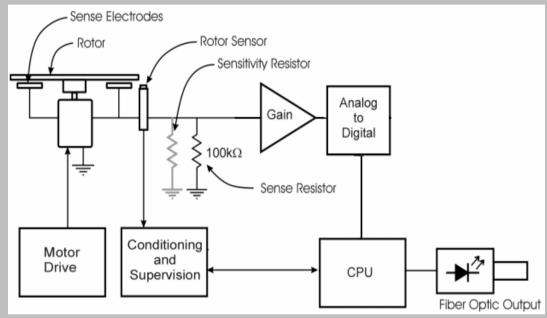
# The Carnegie Curve



3rd ISWI Workshop - 8-12/10/2012 - Quito - Ecuador

# **EFM Measurements**





# **EFM Installation at PLO**





3rd ISWI Workshop – 8-12/10/2012 - Quito - Ecuador

# EFM Installation at ICA

#### Fair weather measurements

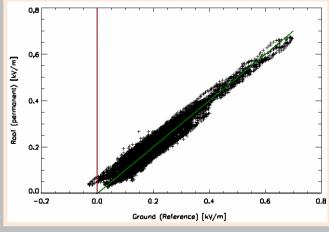


Final location



EF signal acquisition

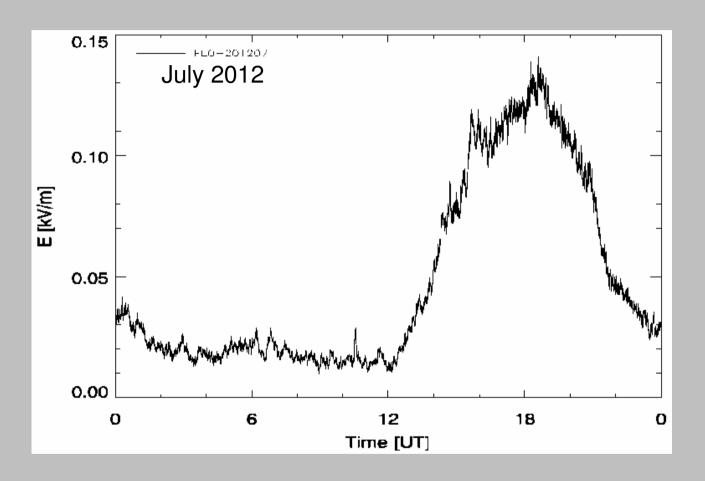




Calibration procedure



## EFM Measurements at PLO



Good local conditions to observe the fair weather atmospheric electric field → 80-90% usable days within a single month of observation

# **EFM Network**



- -PLO and ICA (2011) CONIDA
- Lima and Ancon (10/2012) IGP
- Argentina/CASLEO (since 06/2008)
- 4 more EF sensors in 2013

# COBRAT PROJECT (PI: J.-B. Renard, CNRS-LPC2E, France) 1st Phase approved (06/2012)

- "Long and short duration balloon flights for the study of high energy phenomena and their consequence for stratospheric chemistry".
- Consortium of 34 scientists from 9 countries. ACATMOS and LEONA for Brazilian participation.
- Brazilian experiments: Lightning and TLE imaging, Electric Field Measurements, on-board High-Energy (Solar and Cosmic) particles, Ground-based support.
- Instruments for TGF and TLE detection, imaging, electric and magnetic field measurements, NOx and aerosols concentration measurements.
- Ground-based support by imaging and radio systems
- TARANIS and the ISS European instrument ASIM (TLE, TGF): 2016





# Thermosphere/Ionosphere Sprite halo Sprite halo Sprite streamers Stratosphere Troposphere Troposphere CG

#### **ACATMOS** Team

The Atmospheric Electrodynamical Coupling – ACATMOS team

- Brazil:
- •National Institute for Space Research INPE
- •Instituto Tecnológico SIMEPAR
- •Center for Radioastronomy and Astrophysics Mackenzie CRAAM
- •Lebedev Physical Institute LPI/Russia via CRAAM collaboration
- •Federal University of Mato Grosso do Sul UFMS
- •University of São Paulo USP

#### -Argentina:

- Observatorio Astronómico, Universidad Nacional de Córdoba – UNC
- Consejo de Investigaciones Científicas y Técnicas – CONICET

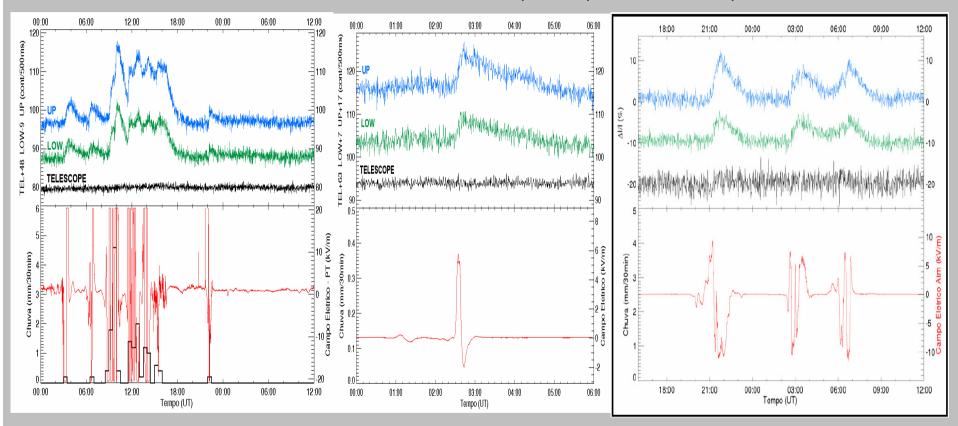
#### -Peru:

•Jicamarca Radio Observatory – JRO, CONIDA



# Short-term modulation of 2nd CR - Atmospheric Phenomena

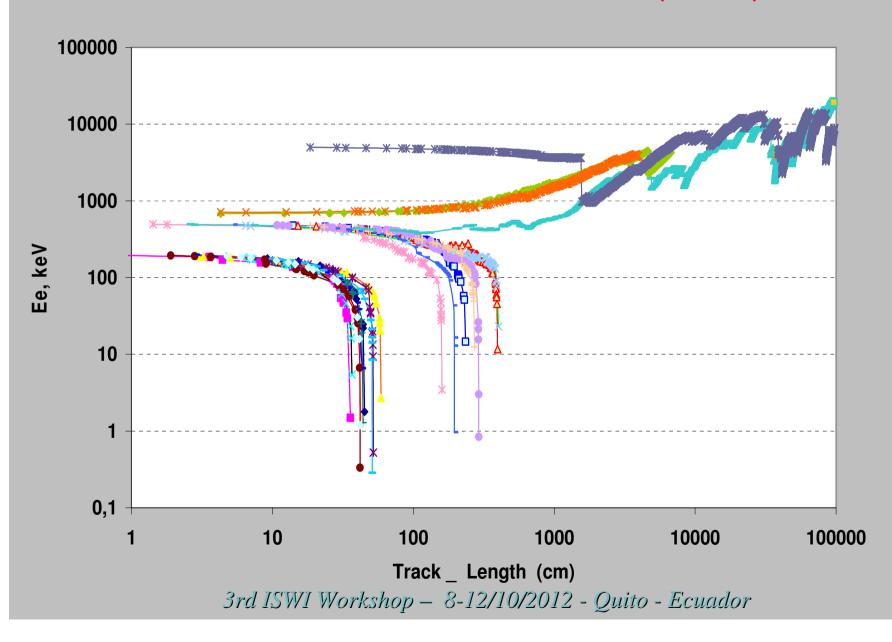
Flux of secondaries CR is corrected for atmopsheric pressure and temperature



RC increase events (LOW, UP) can be observed in association with rain precipitation.

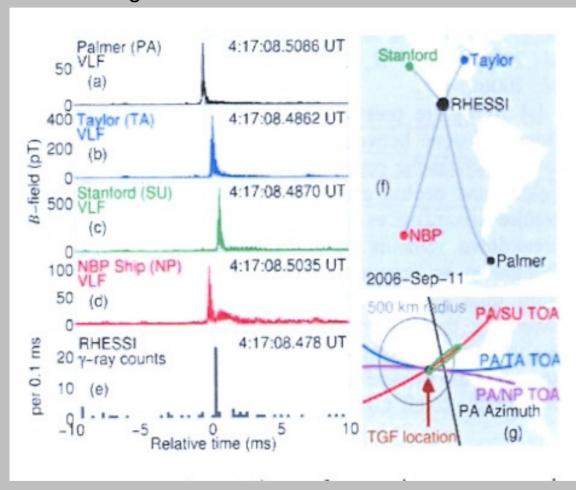
All RC increase events (LOW, UP and TEL) are associated with atmospheric electric field variations.

**GEANT4** simulation E=0.4 kV/cm (40 kV/m)



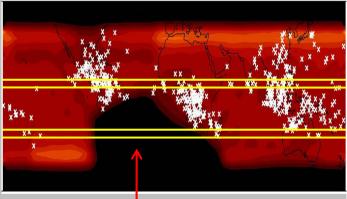
#### TERRESTRIAL GAMMA-RAY FLASHES - TGF

#### Combining with VLF instrumentation



So far ~ 1000 TGFs were detected 1 ms or less Photons up to 20 MeV Electrons up 30 MeV Spectra → OTB Np ~ 10<sup>16</sup> Ne ~10<sup>17</sup>

Altitude ~ 15 - 20 km



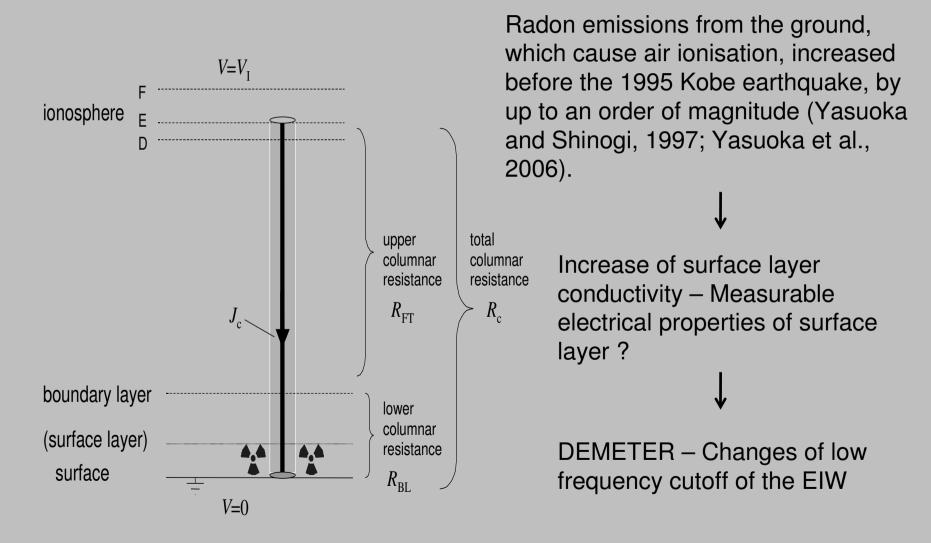
RHESSI - TGFs

No observation because of SAMA

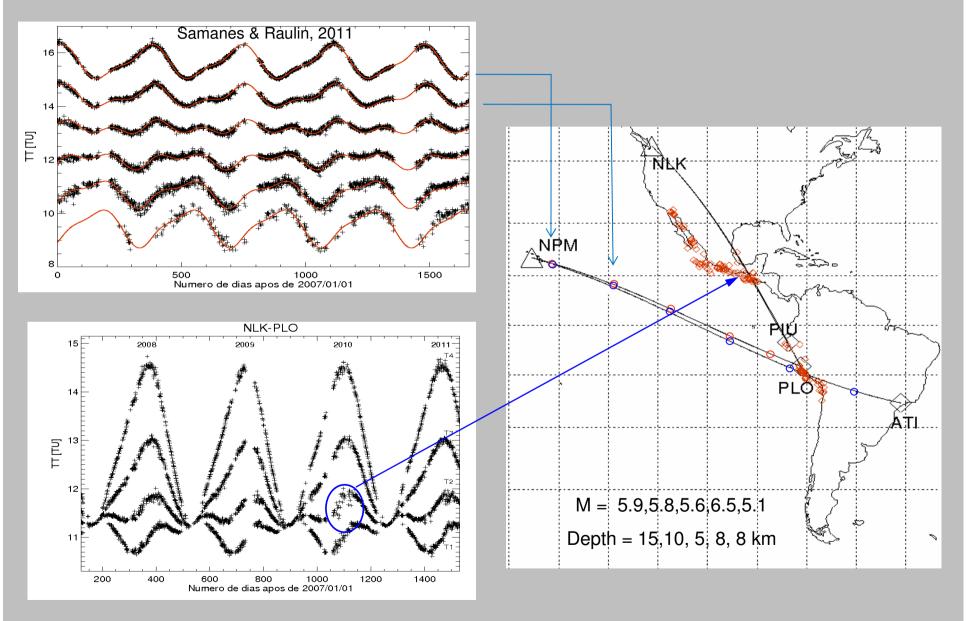
Cohen, M.B., et al. 2010

TGF are associated with lightning activity

# Global circuit changes before land earthquake Concept of Harrison, Aplin, Rycroft, JASTP, **72**, 376, 2010

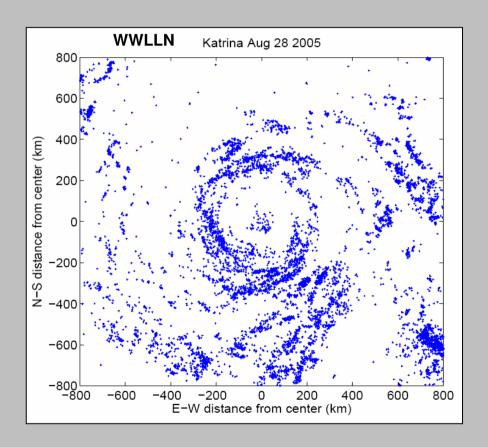


### Seismic – EM effects



## Other Natural Phenomena with Large Societal Impact

Natural disasters: seismic EM effects (volcano, hurricanes, floods etc ...)



(Craig Rodger presentation)

## Other Natural Phenomena with Large Societal Impact

Natural disasters: seismic EM effects (volcano, hurricanes, floods etc ...)

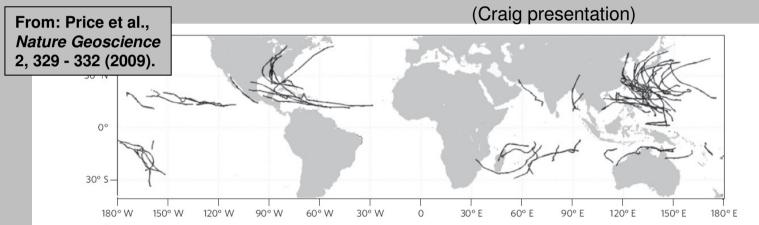


Figure 1 | Global distribution and paths of the 58 hurricanes used in this study. The category-4 and category-5 hurricanes included in this study are: 2005—Adeline, Bertie, Dennis, Emily, Haitang, Ingrid, Katrina, Kenneth, Khanun, Kirogi, Longwang, Mawar, Meena, Nabi, Nancy, Nesat, Olaf, Percy, Rita, Sonca, Talim, Wilma; 2006—Bondo, Carina, Chanchu, Chebi, Cimaron, Daniel, Durian, Ewiniar, Floyd, Glenda, Ioke, John, Larry, Mala, Monica, Saomai, Shanshan, Xangsane, Xavier, Yagi; 2007—Dean, Dora, Favio, Felix, Flossie, Gonu, Indlala, Kajiki, Krosa, Man-yi, Nari, Sepat, Sidr, Usagi, Wipha and Yutu.

Looking at 58 intense hurricanes, cyclones and typhoons, they showed that WWLLN lightning data provided about 30 hours of prediction time as to the future intensity of the storm system.

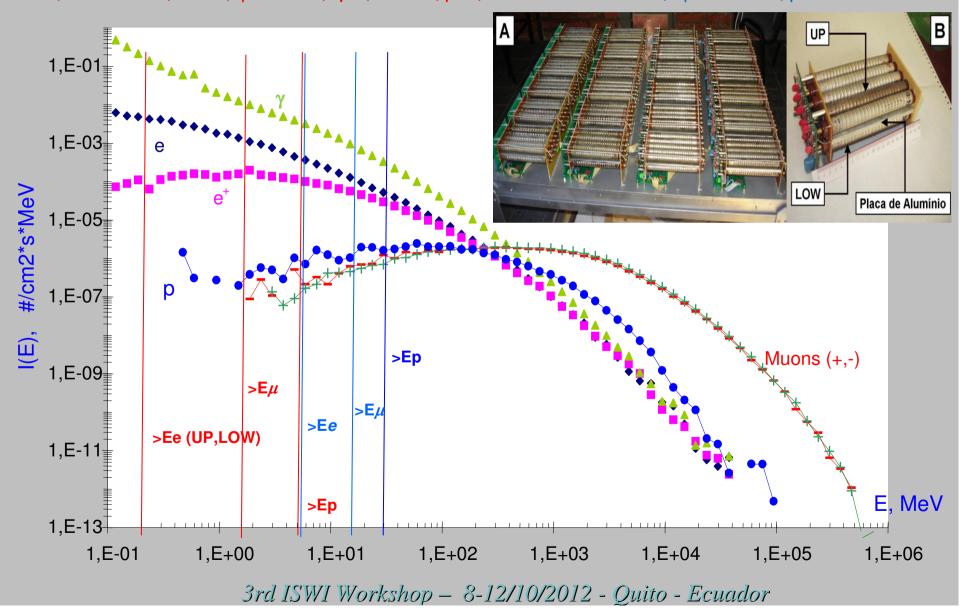
Satellite systems provide a great ability to <u>track</u> and *predict* the storms path, but now we have a possible new route for predicting the changing intensity of these intense storms!

### SUMMARY

- Atmospheric Electric Field sensors have been successfully installed in Peru (PLO, ICA, Lima). Local observational conditions are spectacular, and the use of Fair Weather Data can be optimize.
- Atmospheric Electric Field database, combined with VLF data, will allow to study seasonal and local variability, related to SW or other natural phenomena.
- As in the case of SAVNET deployment and operation, capacity building as the participation and formation of local technicians and students is an important aspect of this new initiative.
- A platform as ISWI supported by UN is very important to continue and improve what has been achieved so far. Such interaction is very relevant for local researchers to look for local funding.

# Differential Energy spectra of CR secondaries p, e<sup>-</sup>, e<sup>+</sup>, $\mu$ +, $\mu$ and $\gamma$ at atmospheric depth X=800 g·cm<sup>-2</sup>

UP, LOW: e> 0,2 MeV, p> 5 MeV,  $\gamma$ >0,02 MeV,  $\mu$ >1,5 MeV TEL: e> 5 MeV, p > 30 MeV,  $\mu$ > 15.5 MeV



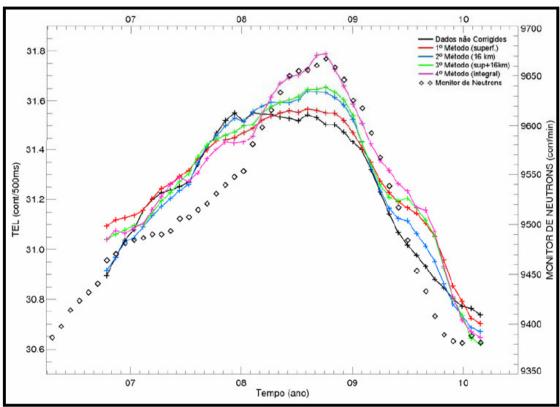


Figura 3.12 – Média corrida de 13 meses das médias mensais da contagem de raios cósmicos sem correção por temperatura (preto), corrigida pela temperatura na superfície (vermelho), pela temperatura na altura de maior produção de partículas secundárias (azul), pela temperatura na superfície e na de maior produção de partículas (verde), e pela integral das temperaturas observadas ao longo da atmosfera (rosa). Os pontos com forma de losangulo representam a média corrida das médias mensais da contagem de raios cósmicos registrada pelo monitor de nêutrons de Moscou. Os dados exibidos entre Abril de 2006 e Julho de 2010.

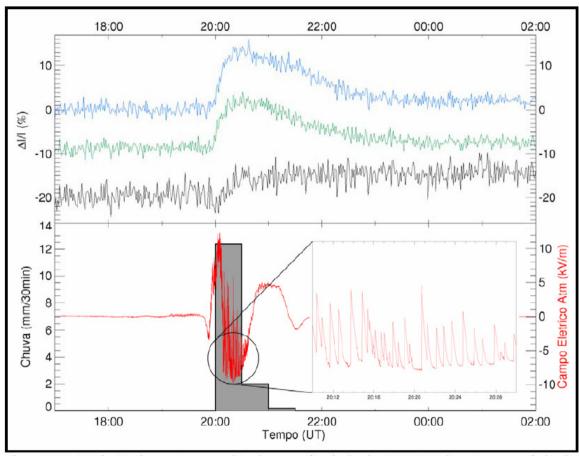


Figura 4.25 – Variação da contagem registrada nos três sinais do CARPET durante um período de ocorrência de chuva (histograma cinza) e de rápidas e intensas variações no campo elétrico atmosférico (curva em vermelho). Dados mostrados entre os dias 10/03/2010 às 17 UT e 24/01/2010 às 02 UT.