

The Sun

Solar physics, Nuclear Physics, Neutrino Physics
Standard Deviation Analysis, Diffusion Entropy Analysis
(space weather phenomena)

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UN/Austria Symposium on

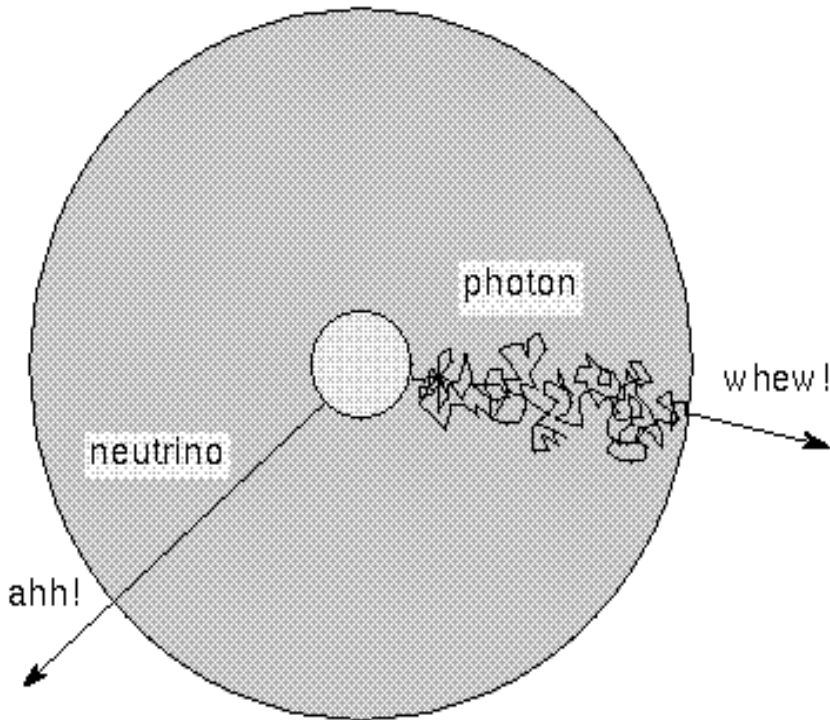
Space Weather Data, Instruments and Models:

Looking Beyond the International Space Weather Initiative (ISWI)

Austrian Academy of Sciences, Institute for Space Research,
Graz, Austria

16-18 September 2013

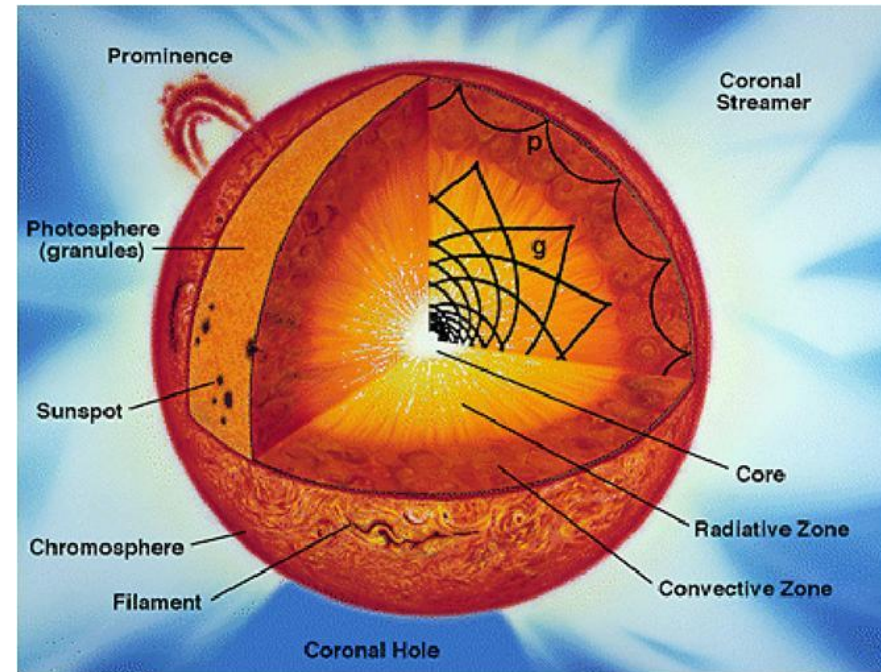
Solar Neutrinos



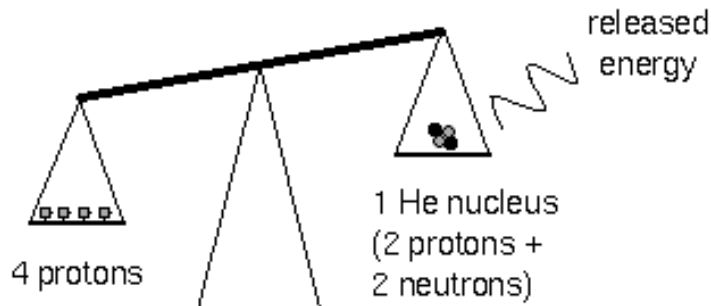
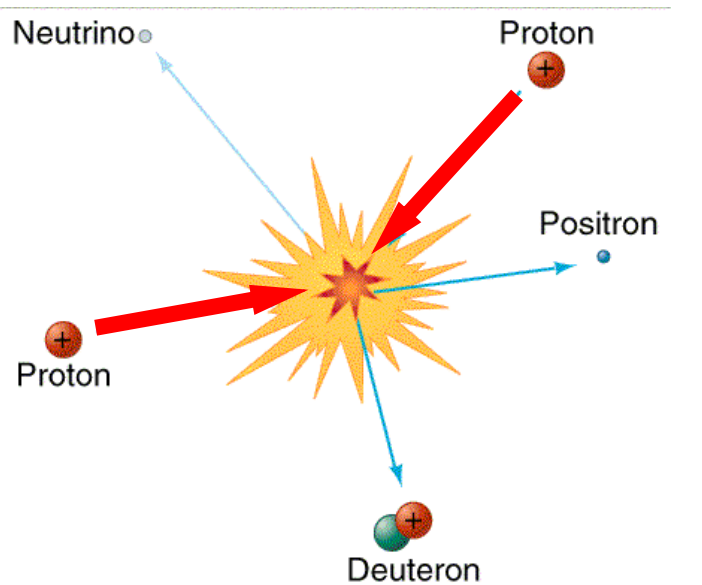
- Gravitationally stabilized solar fusion reactor
- Produced in nuclear fusion reactions
- Exit Sun in about 2 seconds
- Weakly interact with matter
(**neutrino oscillations**)
- Can be detected on Earth
- Flux provides information about solar physics, nuclear physics, and neutrino physics

Standard Solar Model

- Begins with homogeneous composition
- Solar core is **modeled as an ideal fusion plasma**
- Hydrogen burning supplies luminosity and pressure to balance gravity
- Energy is transported by photons
- Chemical composition changes slowly with nuclear reactions
- p-modes and **g-modes?**
- **Magnetic field?**



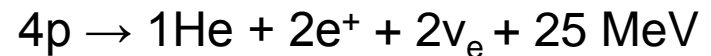
Nuclear Reactions



Some mass is converted into energy ($E=mc^2$)

- Proton- Proton Chain

- overall:



- reactions produce heavier elements

- includes sub-chains involving ${}^7\text{Be}$ and ${}^8\text{B}$

- accounts for 99.6% of the Sun's energy

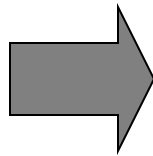
- CNO Cycle

- accounts for 0.4% of the Sun's energy

- dominant in more massive stars

Solar Neutrino Problem

- Observed flux
1/3rd Predicted flux



1. There is some other source of power in the Sun
2. Scientists calculated the reaction rates inaccurately
3. Evidence that neutrinos can change “type” en route to Earth (neutrino oscillations)

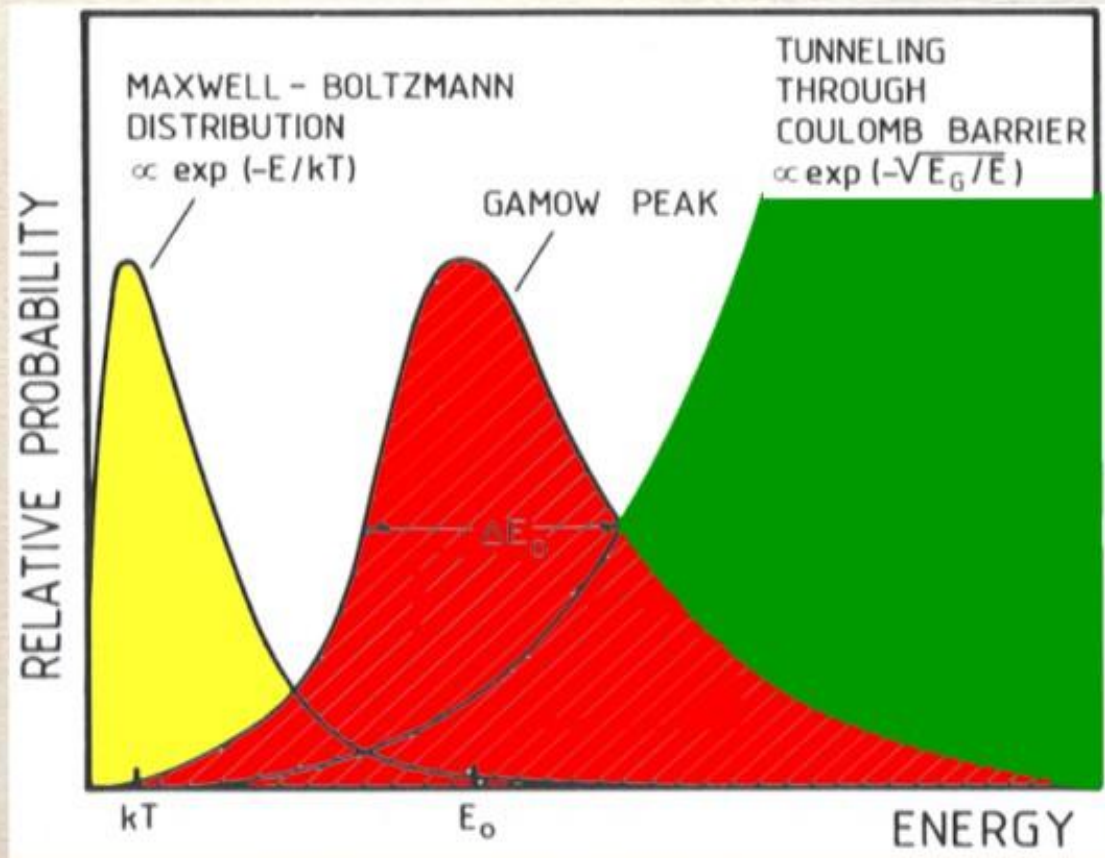
Maxwell-Boltzmann

Maxwell-Boltzmann Distribution

$$\langle \sigma v \rangle = \left(\frac{8}{\mu\pi} \right)^{\frac{1}{2}} \frac{1}{(kT)^{\frac{3}{2}}} \int_0^{\infty} S(E) \exp\left(\frac{-E}{kT} - \frac{b}{\sqrt{E}} \right) dE$$

$$R(r) = \langle \sigma v \rangle N_a N_x$$

$$\Phi = \frac{1}{4\pi R_{ES}^2} \int_0^{R_{sun}} 4\pi r^2 R(r) dr$$



Tsallis

$$p(E) = \begin{cases} N(\delta, kT) \left[1 - \frac{(2\delta)E}{kT} \right]^{\frac{1}{2\delta}}, & E \leq \frac{kT}{2\delta} \\ 0, & \text{otherwise} \end{cases}$$

- Change velocity/energy distribution from Maxwell-Boltzmann to Tsallis
- Vary parameter empirically to match known data

Maxwell-Boltzmann → Tsallis

$$\langle \sigma v \rangle = \left(\frac{8}{\mu\pi} \right)^{\frac{1}{2}} \frac{1}{(kT)^{\frac{3}{2}}} \int_0^{\infty} S(E) \exp\left(\frac{-E}{kT} - \frac{b}{\sqrt{E}} \right) dE$$

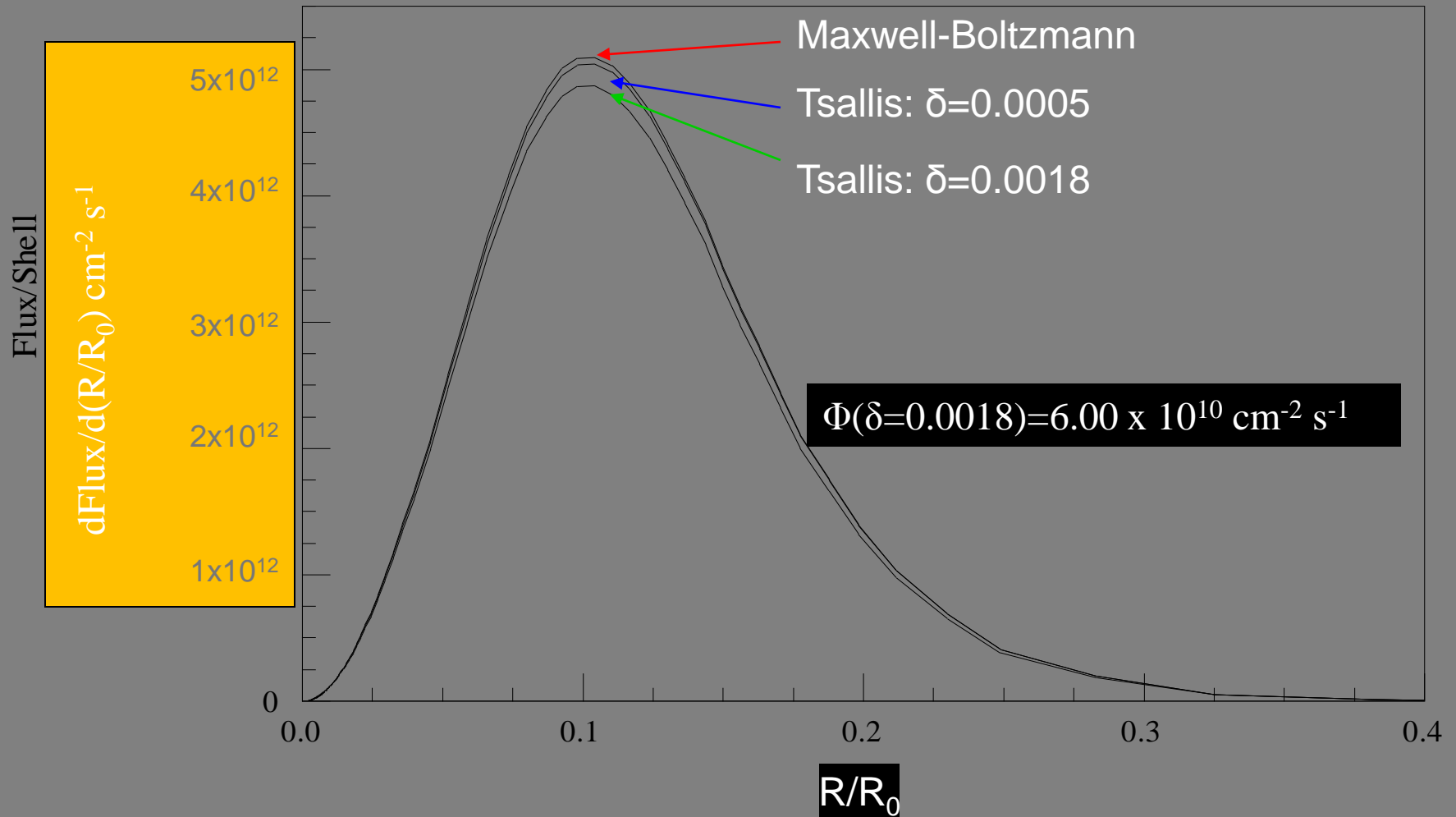
$$\langle \sigma v \rangle = \left(\frac{8}{\mu\pi} \right)^{\frac{1}{2}} \frac{1}{(kT)^{\frac{3}{2}}} \int_0^{\infty} S(E) \exp\left(-\delta \left(\frac{E}{kT} \right)^2 - \frac{E}{kT} - \frac{b}{\sqrt{E}} \right) dE$$

$$R = \langle \sigma v \rangle N_a N_x$$

$$\Phi = \frac{1}{4\pi R_{ES}^2} \int_0^{R_{sun}} 4\pi r^2 R(r) dr$$

PP- Maxwell-Boltzmann vs. Tsallis

Neutrino Production as a function of Solar Radius



Issues to Explore

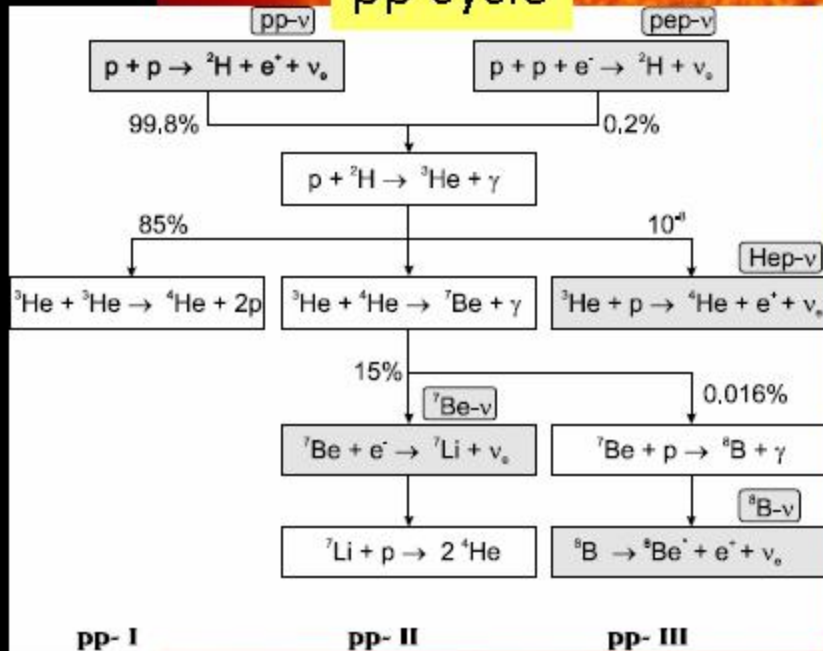
- Find flux for neutrinos from all reactions using Tsallis distribution
- Vary parameter (δ) so that predicted data match observed data
- Implies that the Sun's core is **not an ideal plasma**
- **Magnetic field, g-modes**

Solar Neutrinos

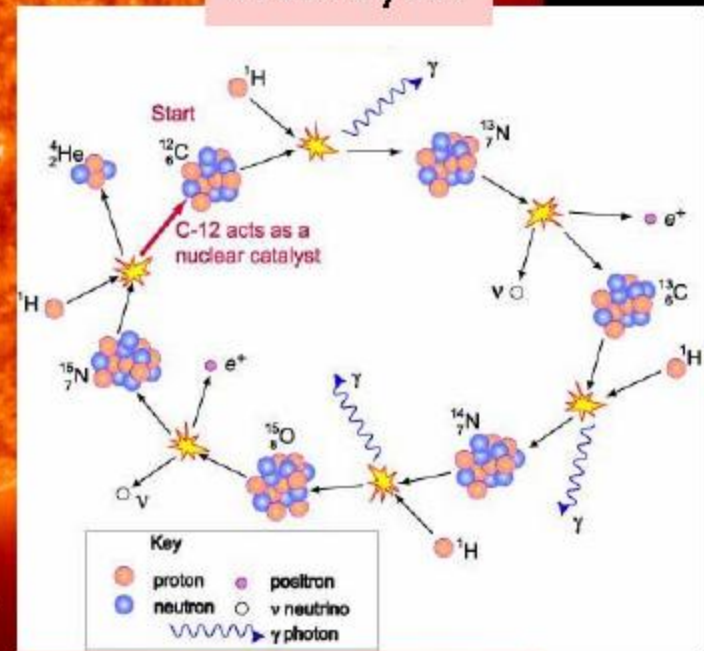


$6.5 \cdot 10^{10} \nu_e/\text{cm}^2\text{s}$

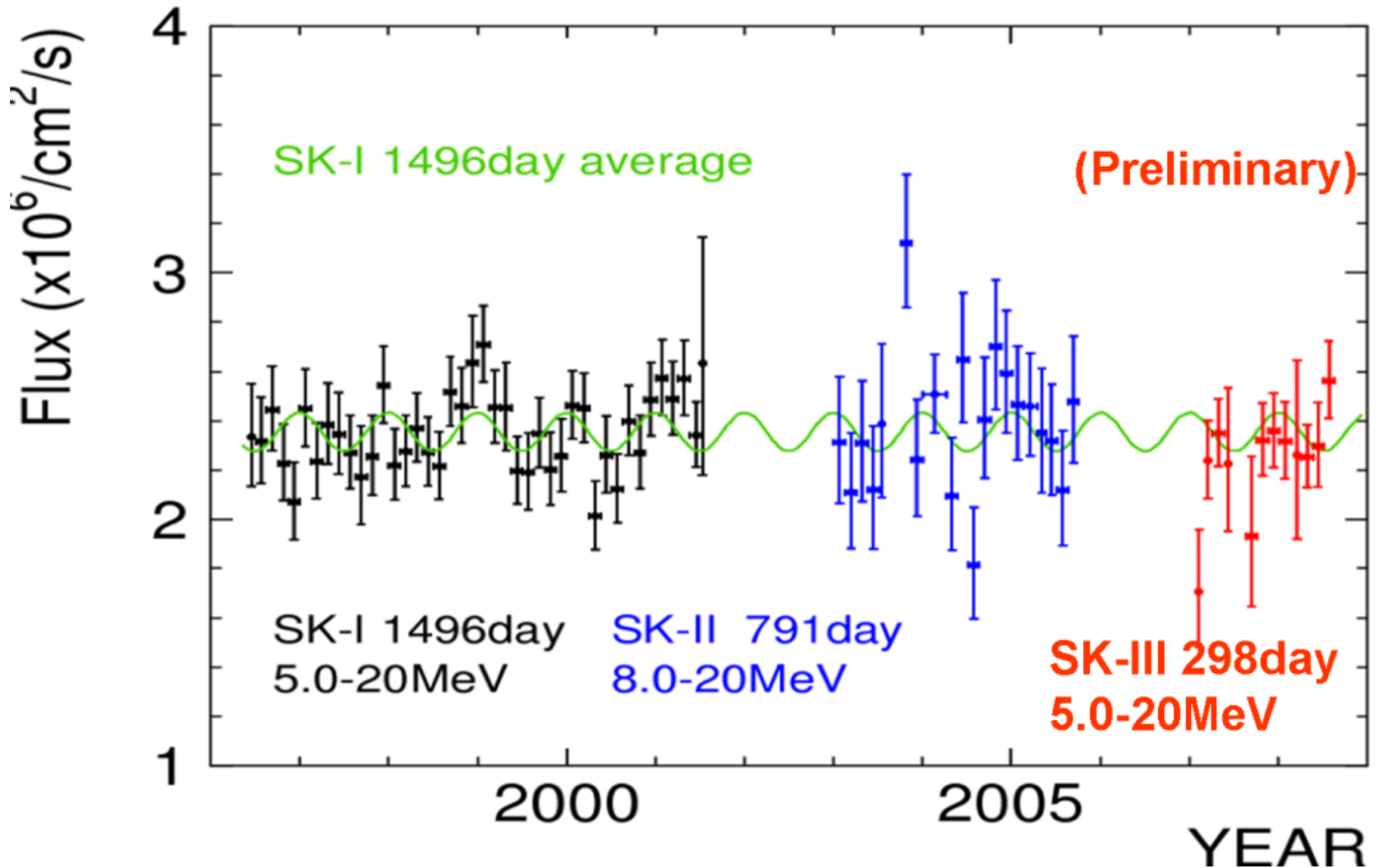
pp cycle



CNO cycle

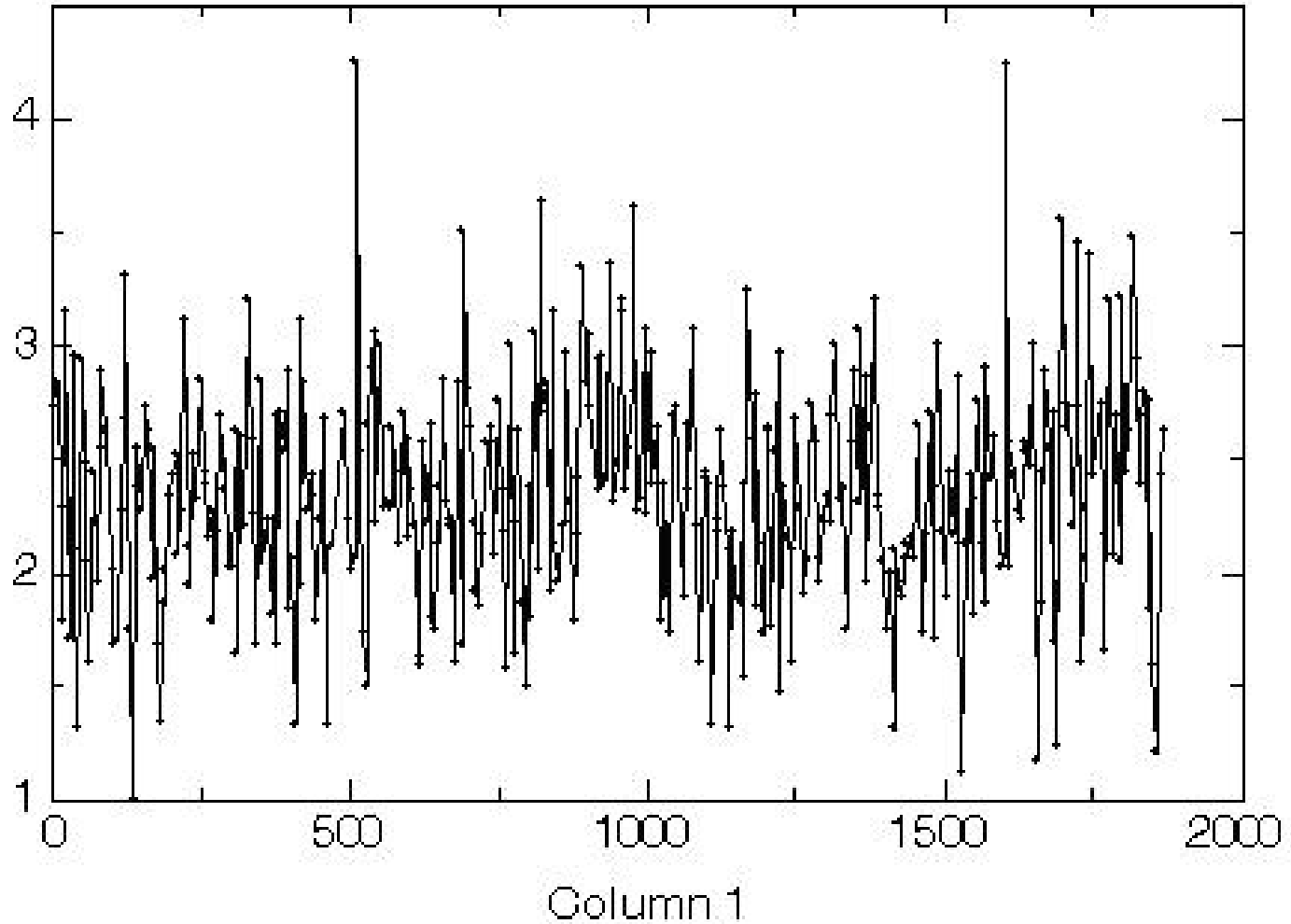


SuperKamiokande I, II, (III)



Time Series

FluxVariation.txt





Mapping Time Series to Diffusion Process

✦ By **summing** the terms of a time series we get a trajectory and the trajectory can be used to generate a diffusion process.

Let us consider a time series $\{\xi_i\}$ of N data: $\xi_1, \xi_2, \dots, \xi_{N-1}, \xi_N$.
For any given time t , $1 \leq t \leq N$, we can find $N-t+1$ sub-sequences :

$$\xi_i^{(s)} \equiv \xi_{i+s} \quad s=0,1,2,\dots, N-t$$

✓ For any of these sub-sequences we can build up diffusion trajectory, defined by the position:

$$x^{(s)}(t) = \sum_{i=1}^t \xi_i^{(s)} = \sum_{i=1}^t \xi_{i+s}$$

Fick's Laws and Brownian motion

$$\vec{J} = -\underline{D} \cdot \vec{\nabla} c$$

1. Fick's Law

$$\frac{dc}{dt} = -\underline{D} \cdot \vec{\nabla}^2 c$$

2. Fick's Law

✦ **Einstein** unified Fick's continuum formulation of diffusion with the Stochastic theory and obtained the probability distribution

$$p(x, t) = \frac{1}{\sqrt{4\pi Dt}} \exp(-x^2/4Dt)$$

Gaussian distribution

Mean square displacement;

$$\langle x^2(t) \rangle \propto t$$



Anomalous Diffusion

Anomalous diffusion is characterised by

$$\langle x^2(t) \rangle \propto t^\gamma \quad \text{with } \gamma \neq 1$$

Mandelbrot introduced a distribution to describe anomalous diffusion.

$$p(x, t) = \frac{1}{\sqrt{4\pi Dt^\eta}} \exp\left(-\frac{x^2}{4Dt^\eta}\right)$$

$$\text{Second moment } \langle x^2(t) \rangle = 2Dt^\eta$$

For $\eta=1$ the normal **Brownian motion** is recovered. The case $0 < \eta < 1$ corresponds to **Sub diffusion** and $\eta > 1$ corresponds to **Super diffusion**.



Scaling Analysis of Time Series

• A **diffusion process** with **scaling** can be described by the probability density function (*pdf*)

$$p(x, t) = \frac{1}{t^\delta} F\left(\frac{x}{t^\delta}\right)$$

δ is called the *pdf* scaling exponent.

✓ The **Variance Scaling exponent** H of a diffusion process is defined by

$$\Sigma^2(t) \sim t^{2H}$$

If $\langle x(t) \rangle = 0$, the variance, $\Sigma^2(t) = \langle x^2(t) \rangle - \langle x(t) \rangle^2$ coincides with the mean squared displacement. Then

$$\Sigma^2(t) = \langle x^2(t) \rangle = \int_{-\infty}^{\infty} x^2 p(x, t) dx \sim t^{2H}$$



Scaling Analysis of Time Series

Now $p(x, t) = \frac{1}{t^\delta} F\left(\frac{x}{t^\delta}\right)$

$$\langle x^2(t) \rangle = \int_{-\infty}^{\infty} x^2 \frac{1}{t^\delta} F\left(\frac{x}{t^\delta}\right) dy$$

$$= t^{2\delta} \int_{-\infty}^{\infty} y^2 F(y) dy$$

Where $y = x/t^\delta$

If $\int_{-\infty}^{\infty} y^2 F(y) dy = \text{Constant} < \infty,$

$$\langle x^2(t) \rangle \sim t^{2\delta}$$

➤ Thus the *pdf scaling exponent* δ and the *variance scaling exponent* H coincide in all cases with $\int y^2 F(y) dy = \text{constant} < \infty$.

✓ This holds true, for example, in the **normal Brownian motion**.



Diffusion Entropy Analysis

- “Diffusion Entropy Analysis (DEA)” finds out the pdf scaling exponent δ .
- DEA is based on the evaluation of the Shannon entropy of the diffusion process.

Shannon entropy:

$$S(t) = - \int_{-\infty}^{\infty} p(x,t) \ln[p(x,t)] dx$$

Using the scaling pdf we get, $S(t) = A + \delta \ln(t)$

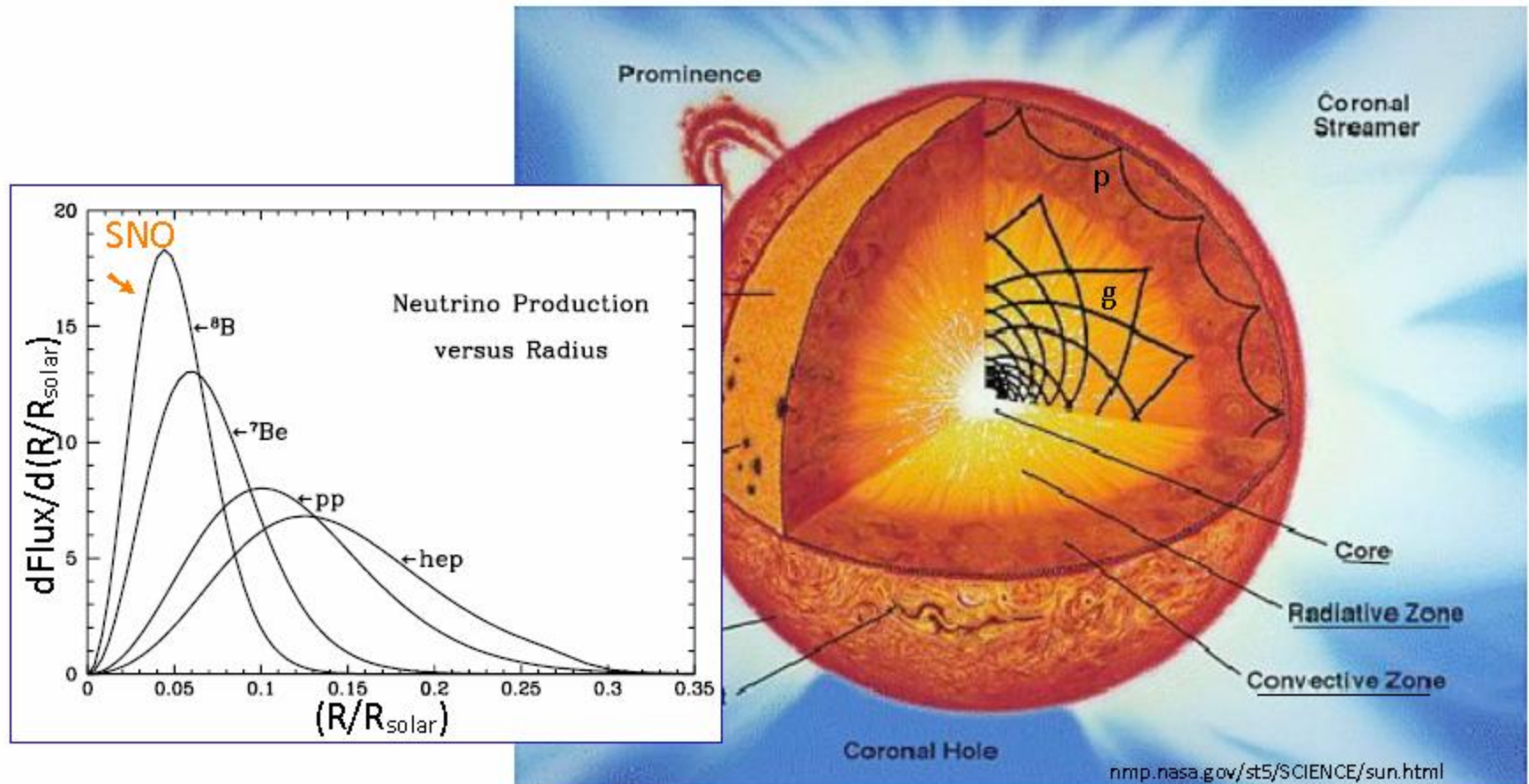
- ✓ The slope of the log-linear plot of $S(t)$ against t gives the pdf scaling exponent δ .



Relation between two scaling exponents

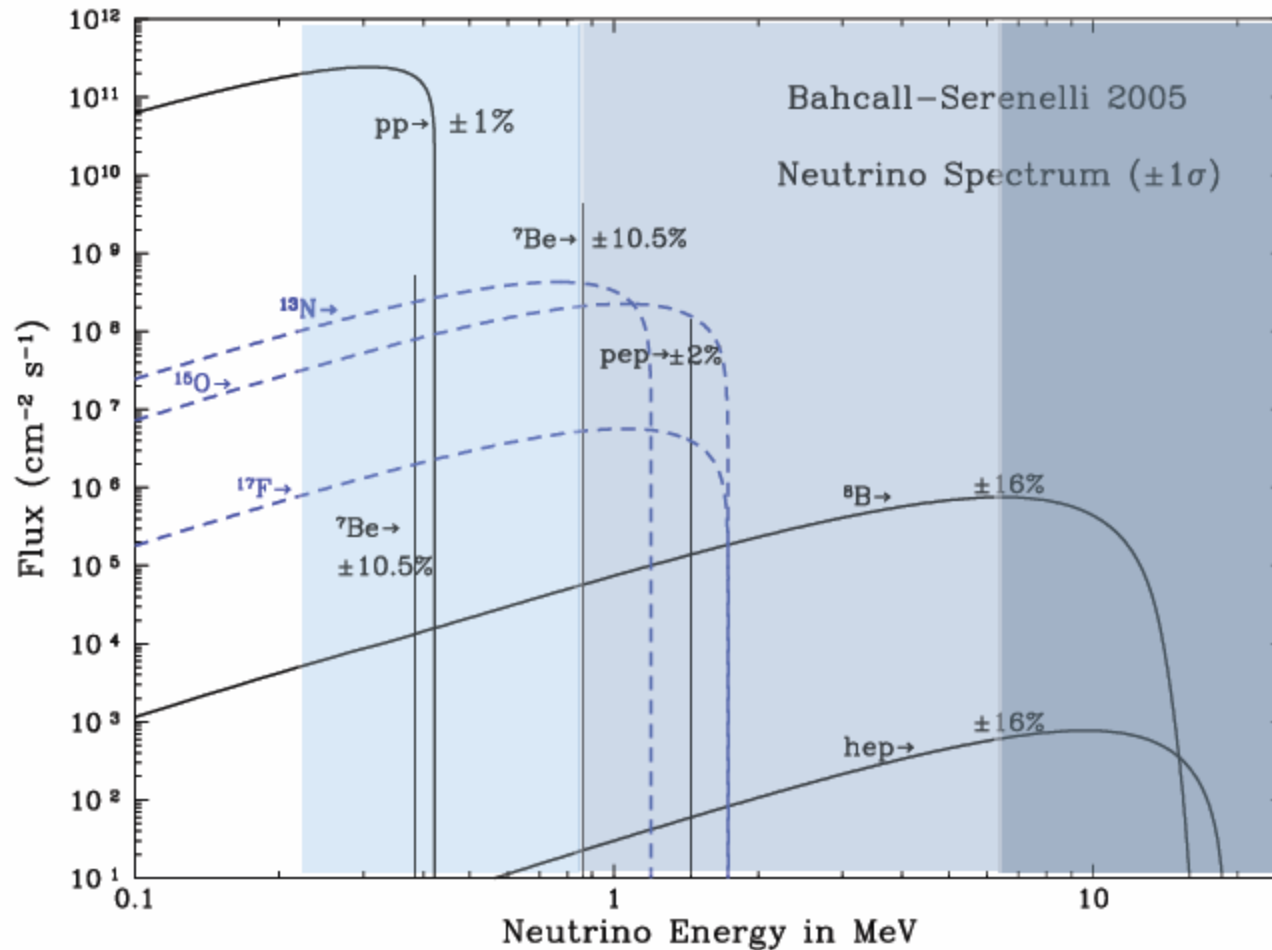
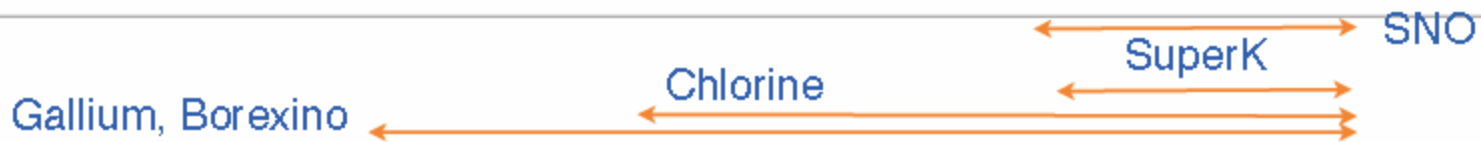
- $H = \delta = 0.5$: Normal diffusion
- $H = \delta \neq 0.5$: Fractional Brownian motion
- $H \neq \delta$: Levy Flight
- $\delta = 1/(3-2H)$: Levy Walk

Standard Solar Model: solar oscillations



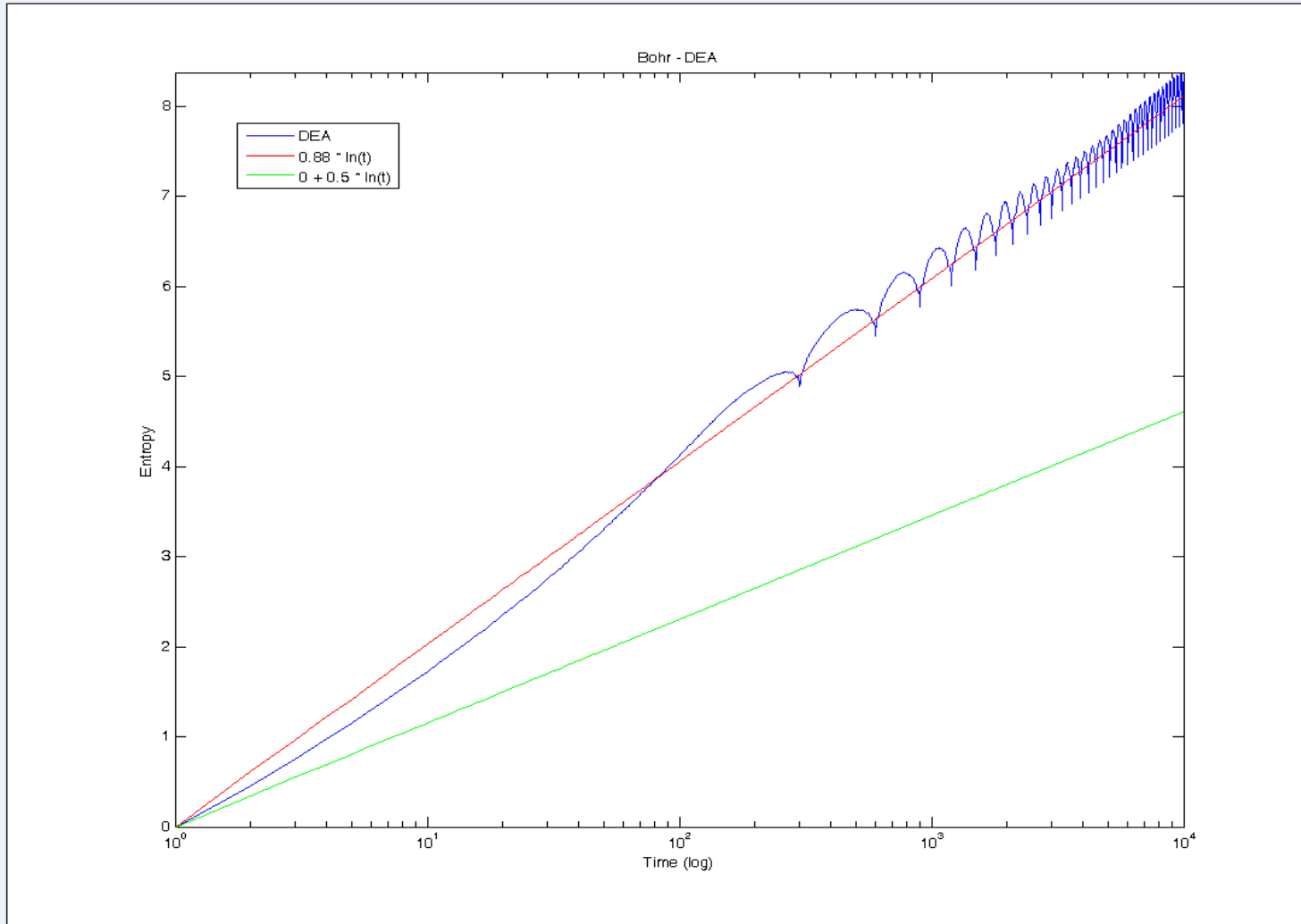
Neutrino production: core
p-mode (acoustic) oscillations: largely convective zone
g-mode (gravity) waves: core and radiative zone

The solar neutrino spectrum



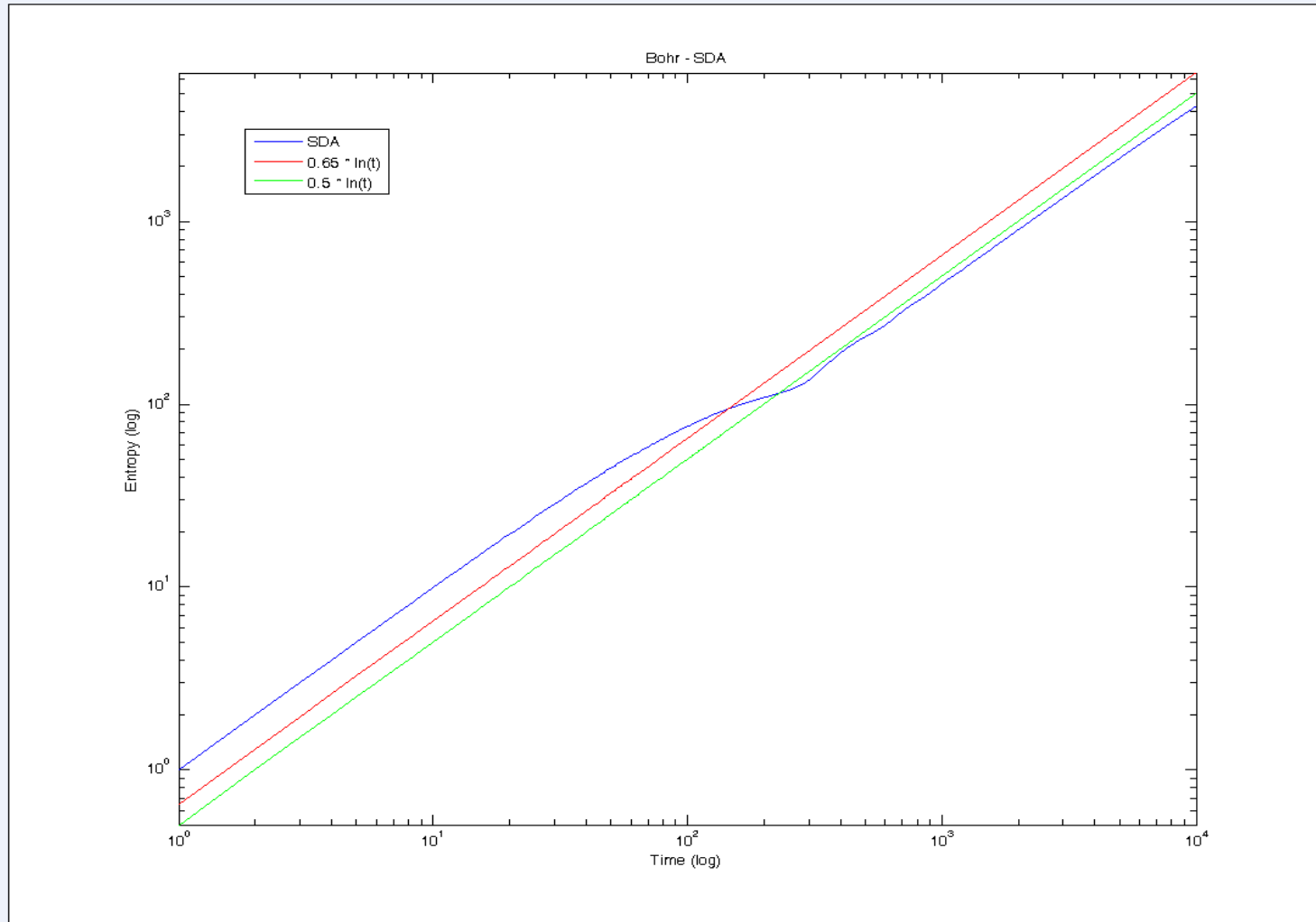
DEA: $\delta=0.88$

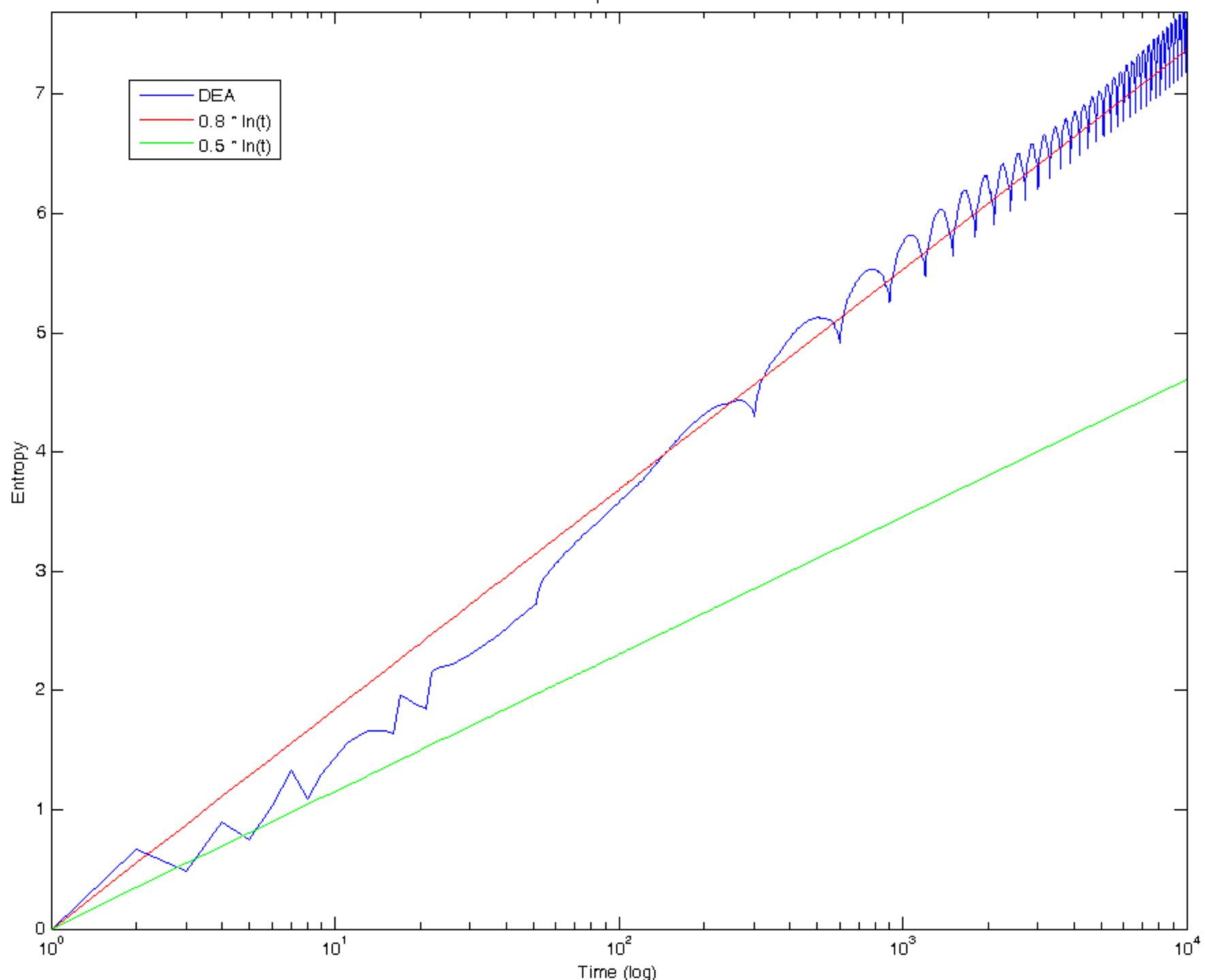
Bohr_DEA.gif - Windows Picture and Fax Viewer

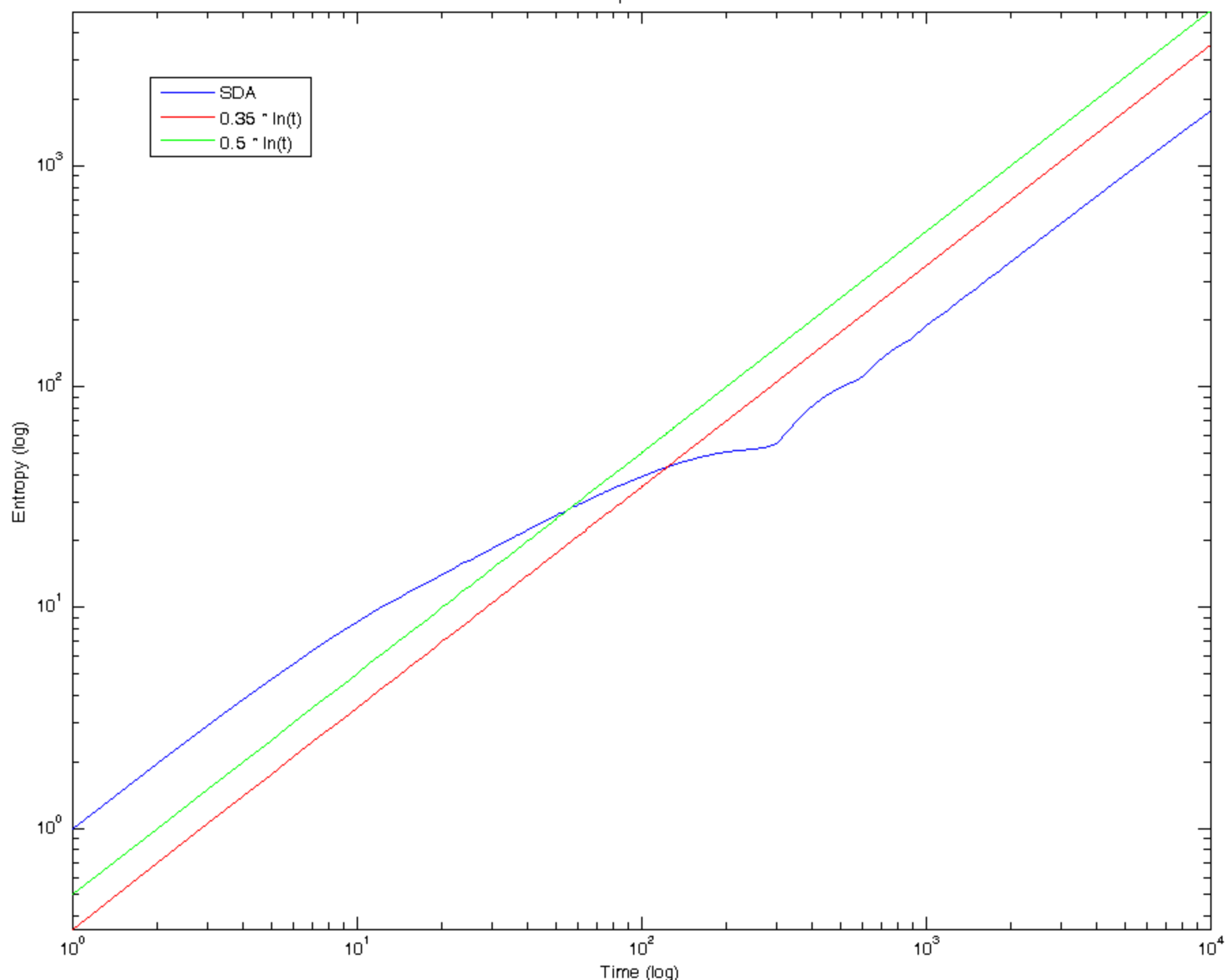


SDA: $H=0.66$

Bohr_SDA.gif - Windows Picture and Fax Viewer







Shannon Entropy rather than Mean Squared Displacement!

(DEA-pdf versus SDA-variance?)

SK solar neutrino signal is not GAUSSian

(TSALLIS, LÉVY, ...?)

SK solar neutrino signal shows periodic modulation

(g-modes, magnetic field, ...?)

Thank you for your attention!

- United Nations Office for Outer Space Affairs
- Centre for Mathematical Sciences India