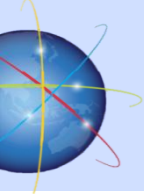


# **The impact of “quiet” space weather on technologies: Research progress on plasma bubbles and induced currents**

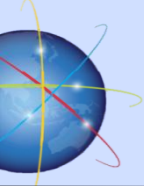


**Brett A. Carter**



# Outline

- **Introduction**
  - Extreme space weather events can be costly
- **What about “daily” space weather?**
  - Equatorial Plasma Bubbles (EPBs), which effects GPS and satellite communications
    - Our recent efforts to produce a day-to-day prediction capability
  - Geomagnetically Induced Currents (GICs) at equatorial latitudes
    - Our analysis of GICs caused by interplanetary shocks
- **Summary and conclusions**



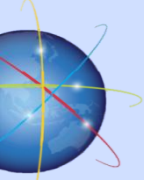
# The cost of extreme space weather

- Severe space weather event scenario, recovery 4-10 years, \$US 1-2 trillion in first year alone<sup>1</sup>
- Tohoku Earthquake, estimated \$US 210 billion<sup>2</sup>
- Hurricane Katrina, 2005, \$US 81-125 billion<sup>1</sup>
- Potential Geostationary Satellite revenue loss, \$US 25 billion per year<sup>1</sup>
- Late 2010-early 2011 Queensland floods, \$AU 5-6 billion<sup>3</sup>

<sup>1</sup> National Research Council. *Severe Space Weather Events--Understanding Societal and Economic Impacts: A Workshop Report*. Washington, DC: The National Academies Press, 2008 (and references therein).

<sup>2</sup> Brookings-Bern Project on Internal Displacement, *The Year that Shook the Rich: A Review of Natural Disasters in 2011*, March 2012, available at: <http://www.refworld.org/docid/4f61a85a2.html> [accessed 10 February 2014]

<sup>3</sup> [http://www.chiefscientist.qld.gov.au/images/documents/chiefscientist/understanding-floods\\_full\\_colour.pdf](http://www.chiefscientist.qld.gov.au/images/documents/chiefscientist/understanding-floods_full_colour.pdf)



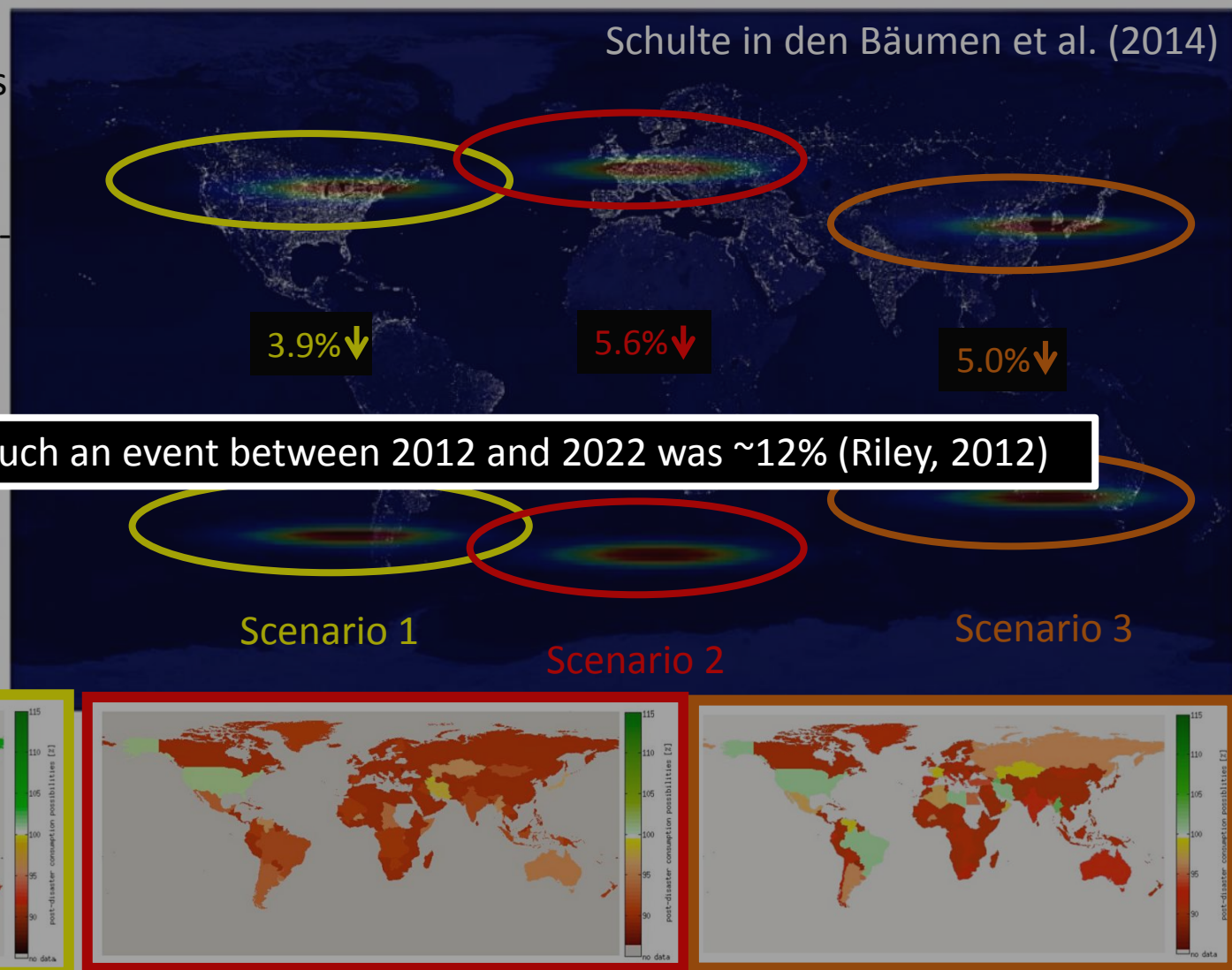
# The cost of extreme space weather

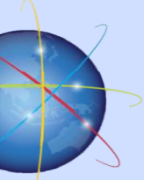
Max of 10% of electricity supply lost over 12 months

Global economic model

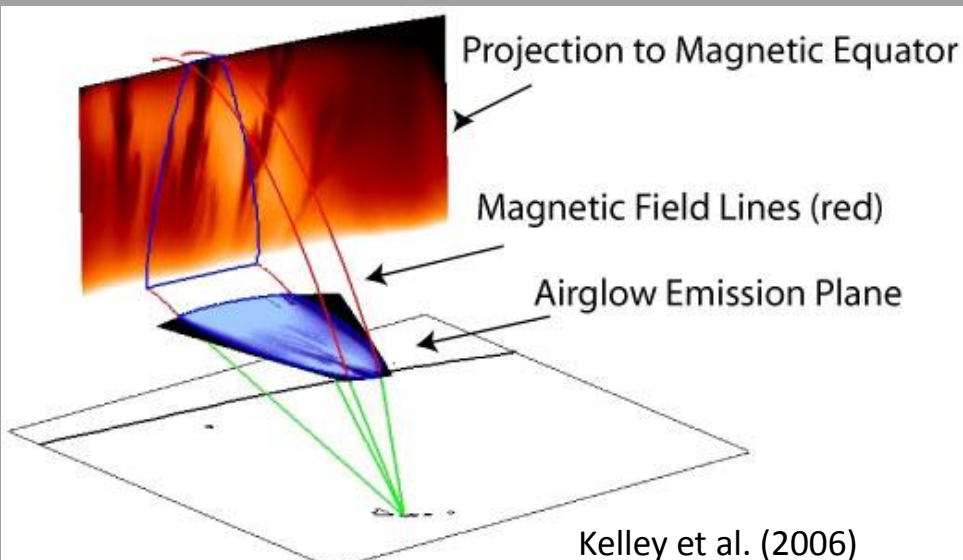
- 187 countries with 25-400 sectors per country
- captures ~99% of

The likelihood of such an event between 2012 and 2022 was ~12% (Riley, 2012)





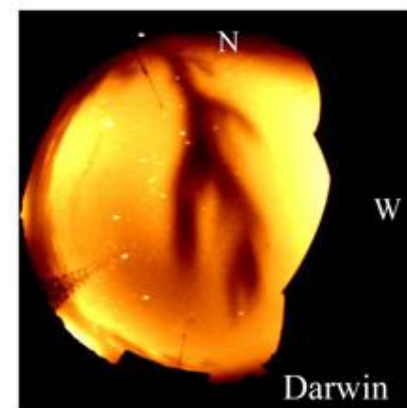
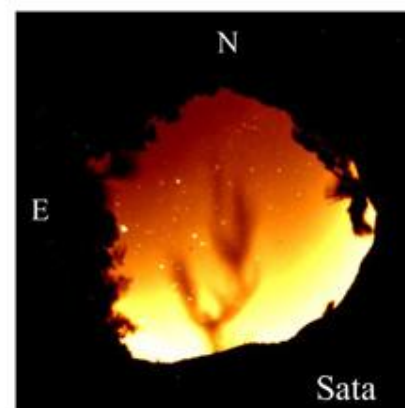
# Equatorial Plasma Bubbles



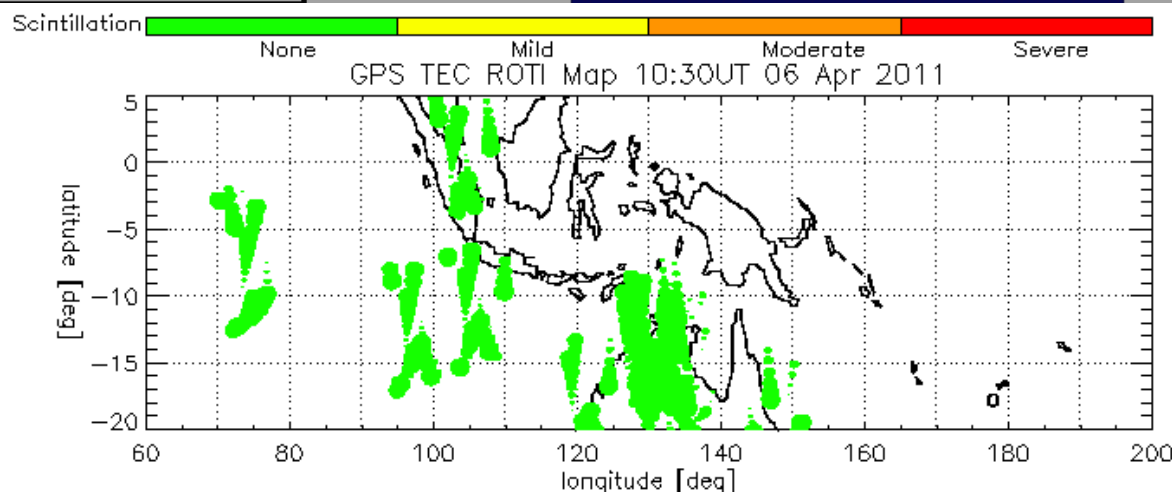
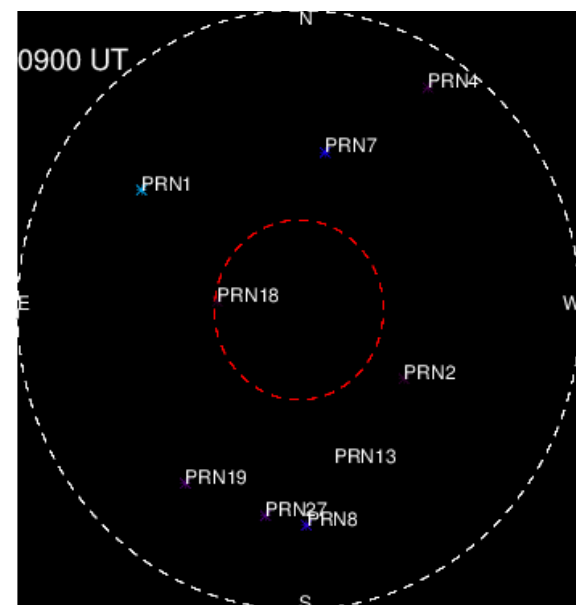
Vanimo PNG, Mar-17-2000

All-sky cameras and numerical modelling

[http://center.stelab.nagoya-u.ac.jp/site1/info\\_e/kagoshima.html](http://center.stelab.nagoya-u.ac.jp/site1/info_e/kagoshima.html)



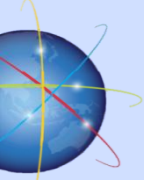
1200



VarSITI / B. A. Carter

Retterer [2008a,b] West - East [km]

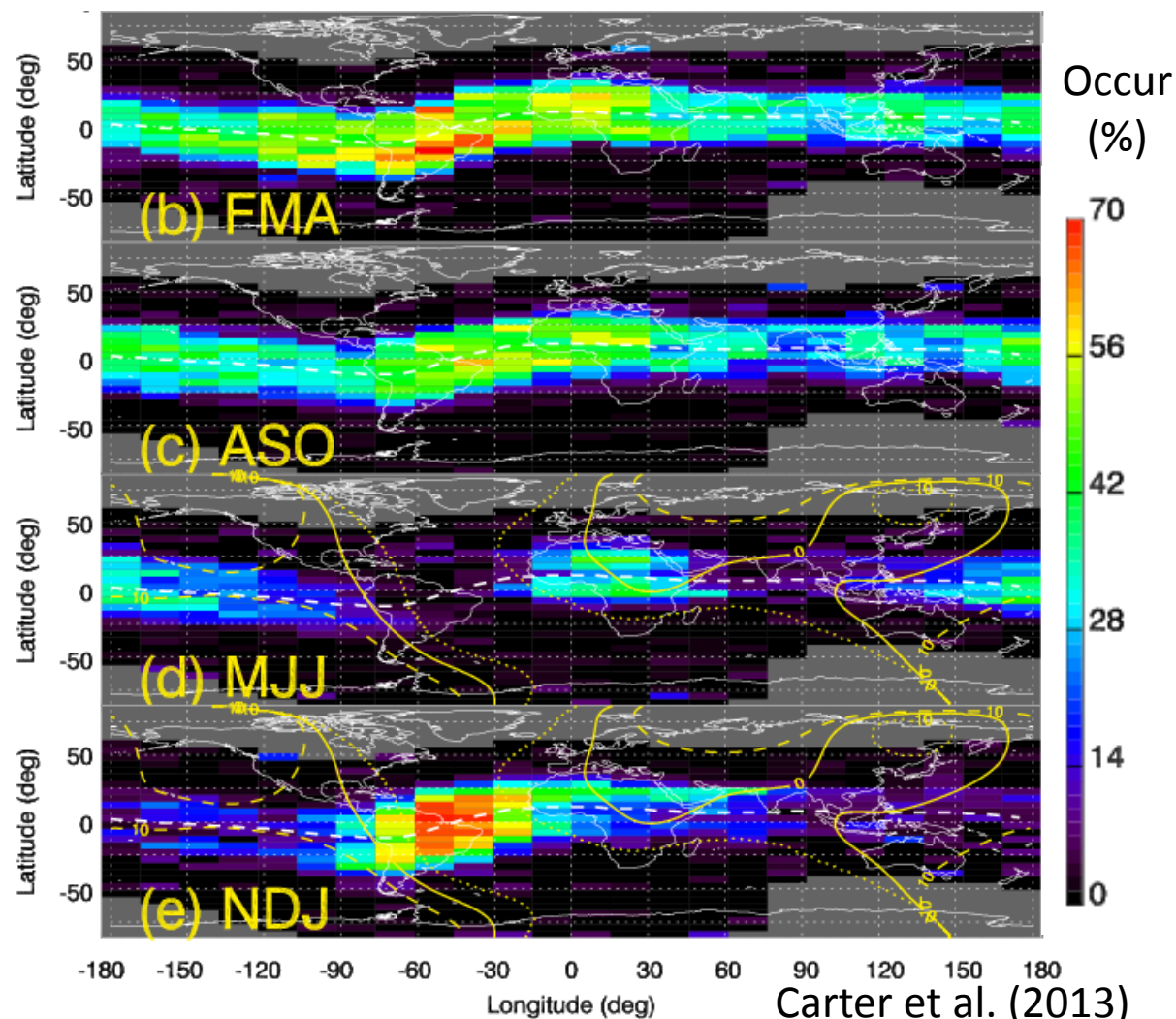




# EPB Climatology: GPS RO data

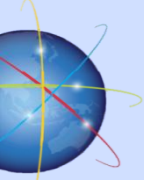
EPBs (and scintillation events on GPS and UHF systems) are most common during the equinoxes across all longitude sectors near the magnetic equator.

The June solstice months see elevated EPB activity in the Australian/Pacific longitude sector, and the December solstice shows increased EPB activity in South America and Africa.

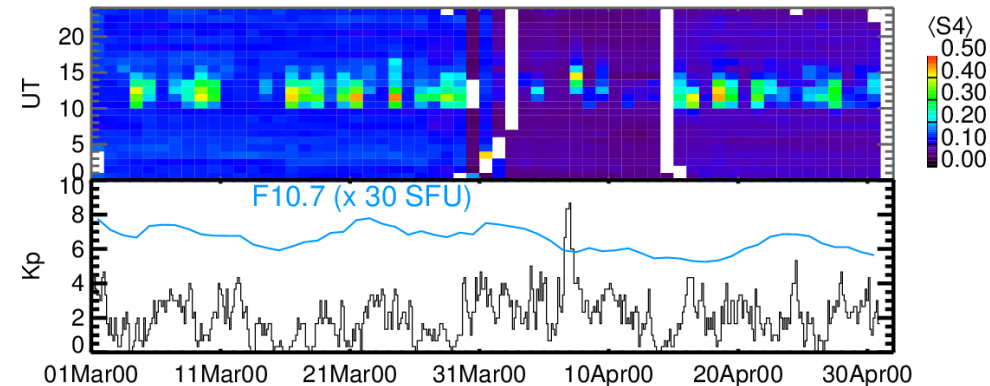
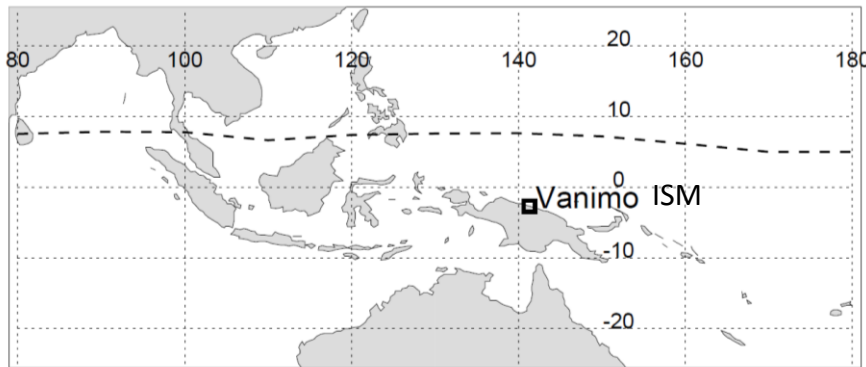




7



# GPS scintillation observations



Carter et al., 2014 [JGR]

- Ionosphere - thermosphere observations along the entire flux tube, as required by the Rayleigh-Taylor linear instability growth rate expression, are not possible/feasible

(Sultan, 1996)

$$\gamma = \frac{\Sigma_P^F}{\Sigma_P^E + \Sigma_P^F} \left( V_p - U_n^P - \frac{gL}{v_{in}^{eff}} \right) \frac{1}{L_n} - R_T$$

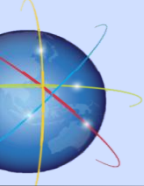
Labels and arrows pointing to terms in the equation:

- Unknown** (points to  $\Sigma_P^F$ )
- Pederson conductivities** (points to  $\Sigma_P^E + \Sigma_P^F$ )
- Upward plasma drift** (points to  $V_p$ )
- Gravity** (points to  $gL$ )
- Upward neutral wind** (points to  $U_n^P$ )
- Ion-neutral collision frequency** (points to  $v_{in}^{eff}$ )
- Recombination rate** (points to  $R_T$ )
- Gradient scale length** (points to  $L_n$ )

**Directly measured/known**

- Therefore, some form of ionosphere-thermosphere modelling is required...





The Thermosphere Ionosphere Electrodynamics General Circulation Model (TIEGCM) is a time-dependent 3D physics-based (i.e. not empirical) numerical simulation of the Earth's thermosphere and ionosphere.

## Inputs:

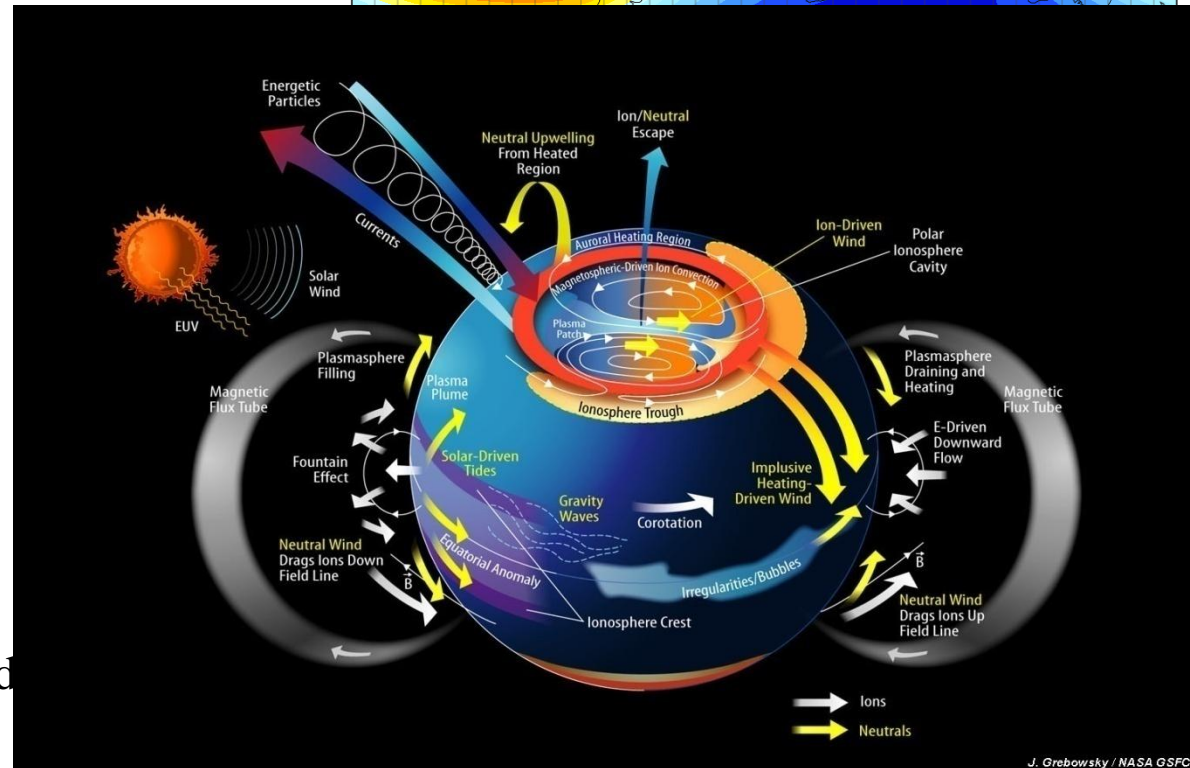
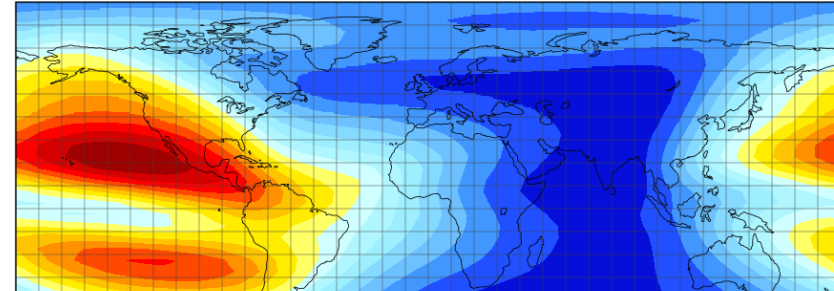
- Solar activity (F10.7 cm flux)
- Geomagnetic activity (Kp index)

## Outputs:

- Electron density
- F layer height
- 3D plasma drift
- Thermospheric density
- 3D neutral winds...
- ...
- Basically, everything that we need

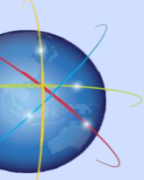
Peak Density of the F2 Layer

Time: 2000-02-04 00:00:00

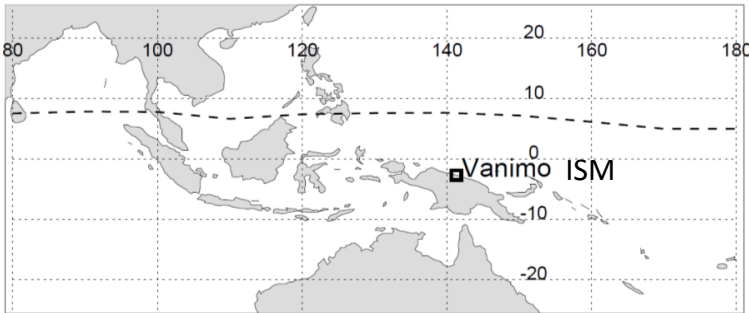


Data Min =  $-2.6 \times 10^4$ , Max =  $2.6 \times 10^4$

J. Grebowsky / NASA GSFC



# TIEGCM: EPB variability

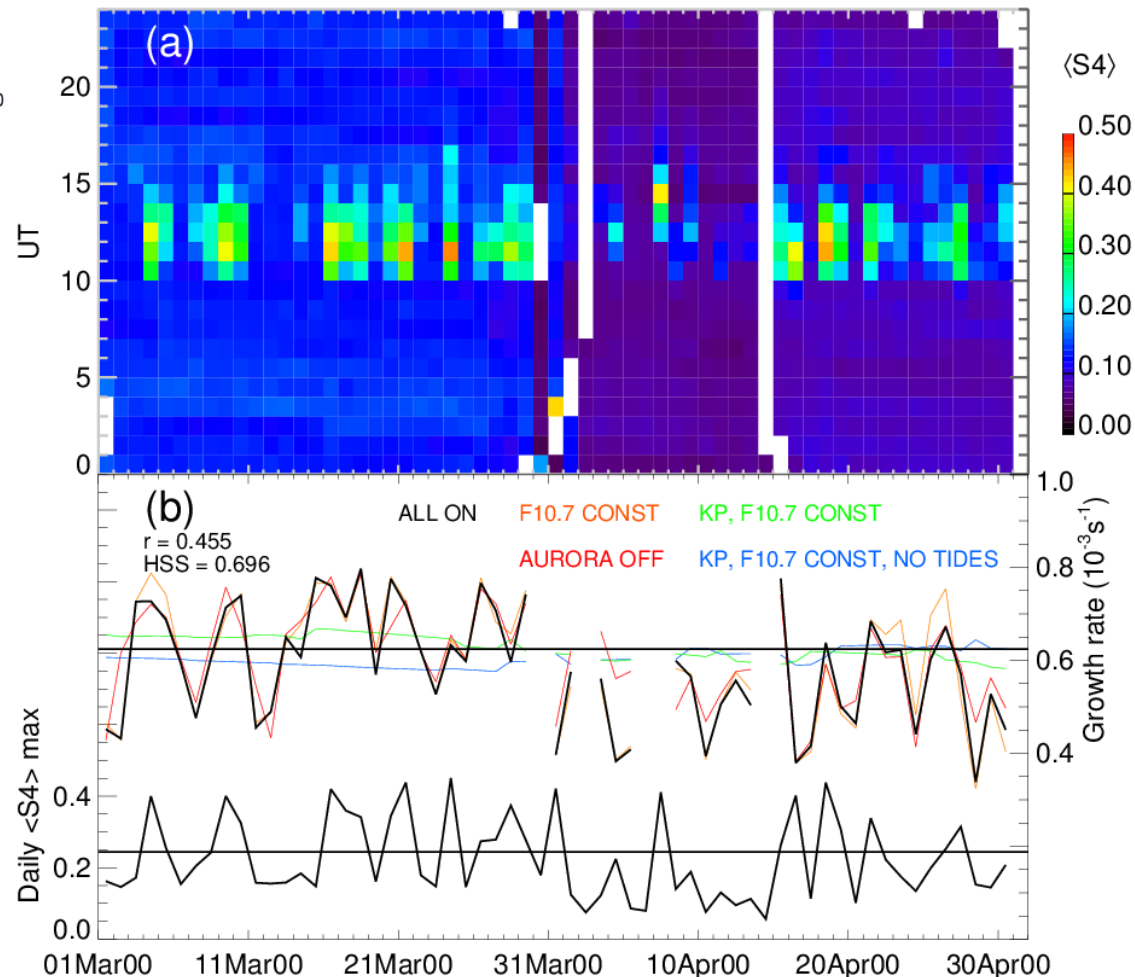


**Observed**

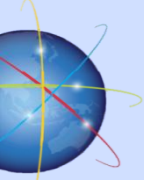
**Modelled**

EPBs	Yes	No
Yes	17	3
No	5	31

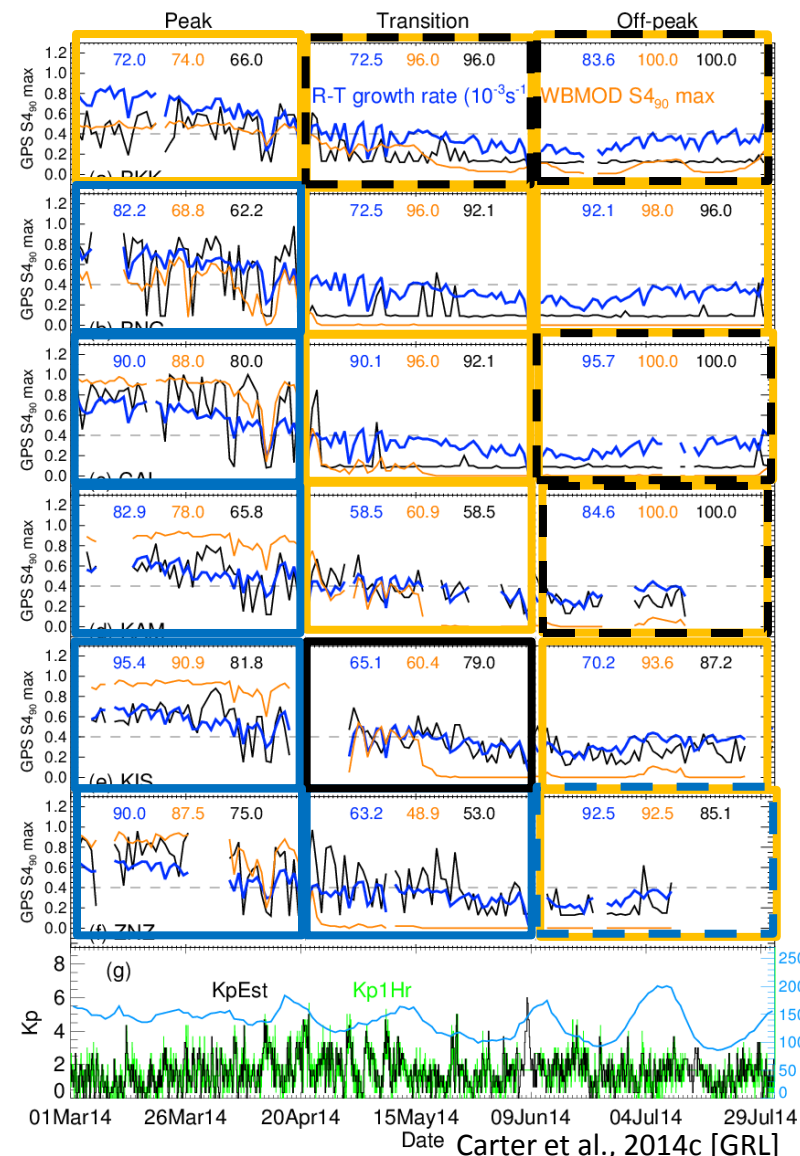
- Physics-based model was found to imitate the observed daily changes
- Kp is dominant source of TIEGCM variability during quiet period



Carter et al., 2014a [JGR]



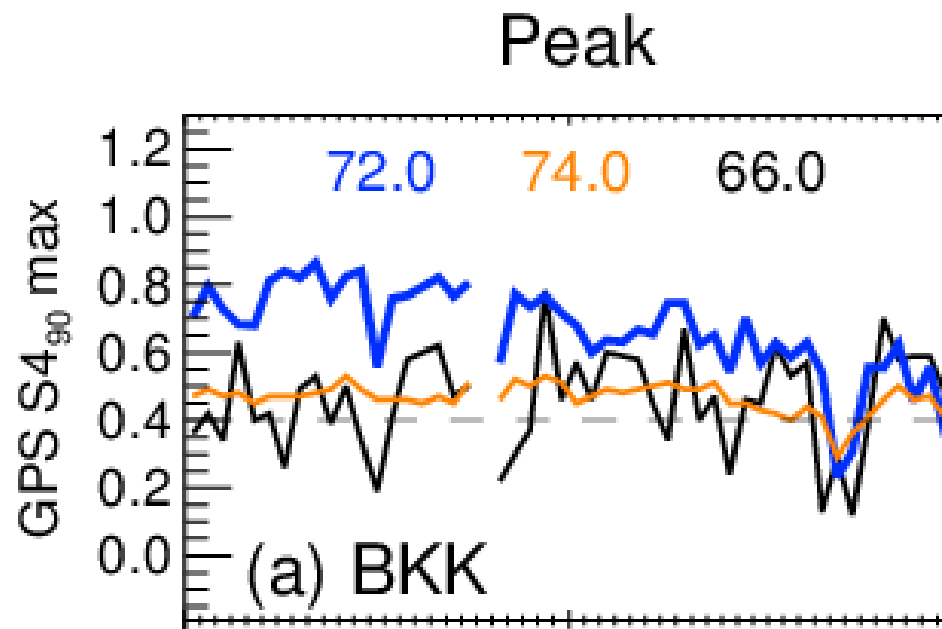
# Scintillation prediction trial: Mar-Jul 2014



## 1-hour Wing Kp predictions:

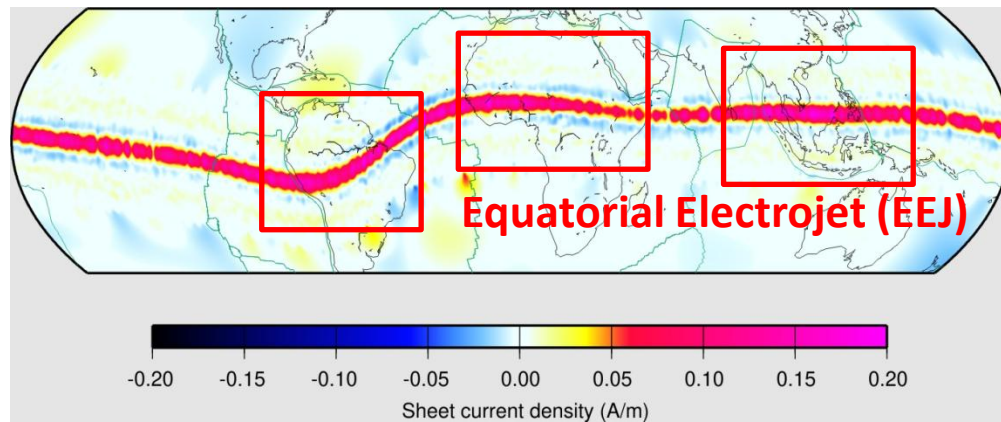
Our technique generally performs best during peak EPB season, closely followed by AFRL's WBMOD (up to 95% for KIS)

During transition and off-peak seasons, either WBMOD or "persistence" forecast performs best



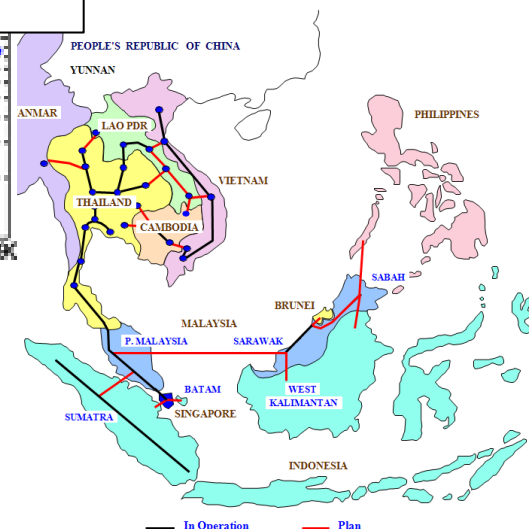
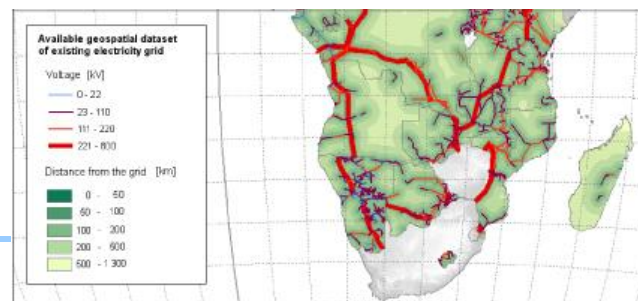
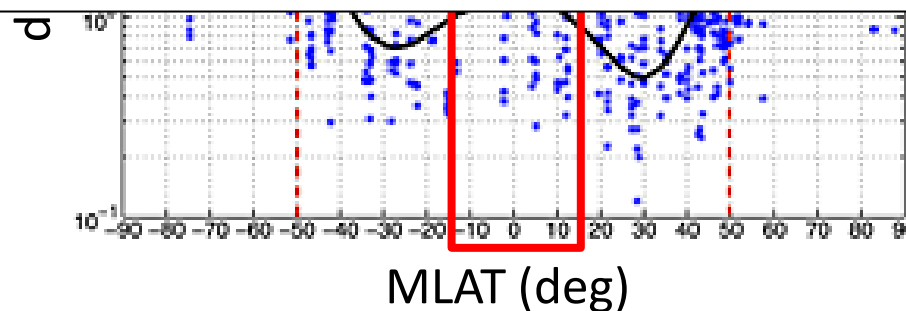
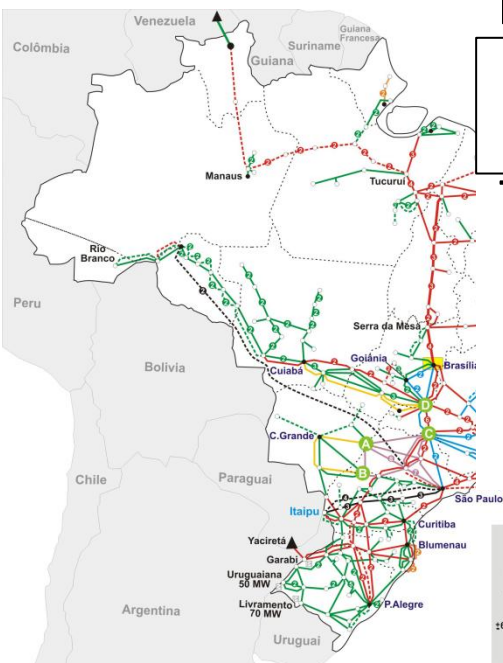


# Space weather effects on technologies – GICs



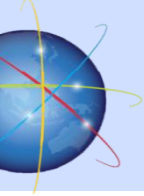
(2013)

EEJ could cause increased vulnerability to GICs  
in equatorial region

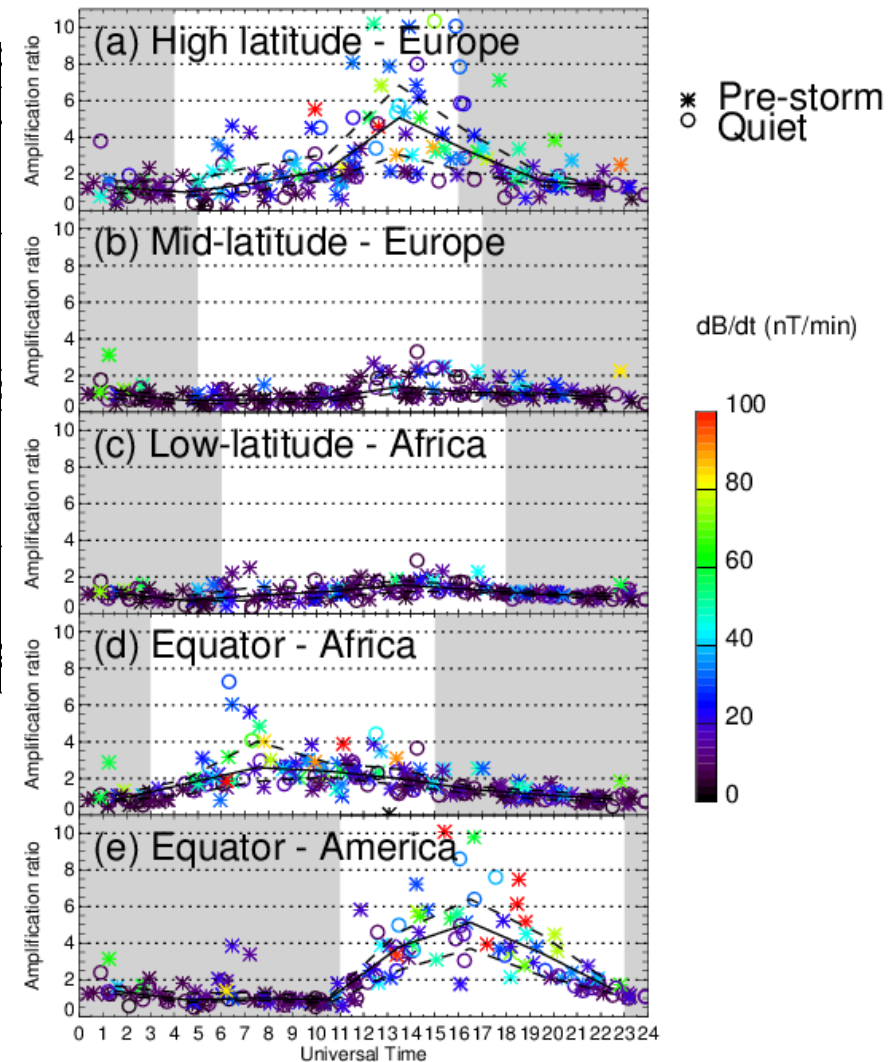
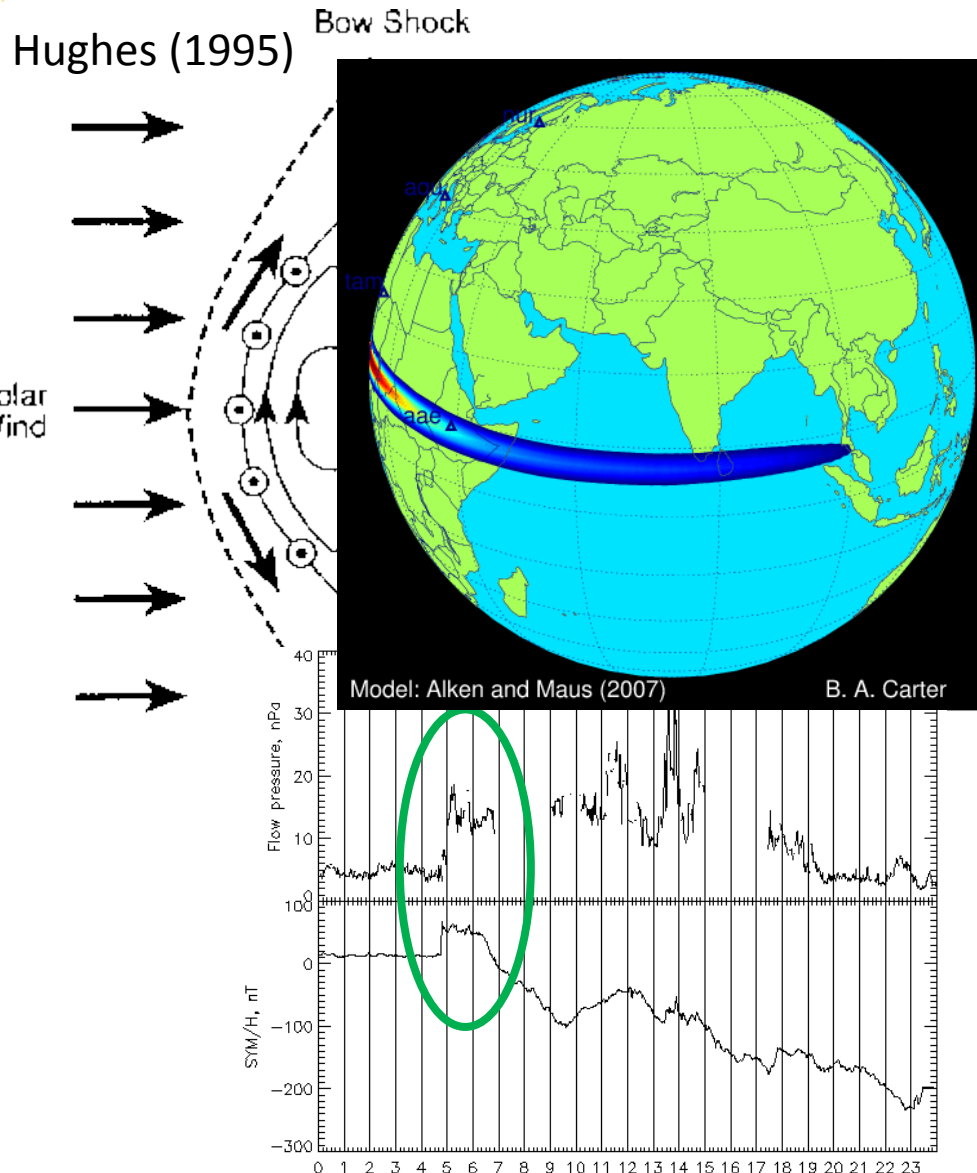


Moldwin et al., AGU Monograph (2016)





# Interplanetary Shocks and GICs



Carter et al., 2015 (GRL)



